

Optical witnesses for quantum correlations in cavity-driven atomic clouds and other challenges

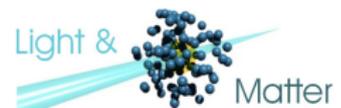
Strasbourg 2026

Philippe W. Courteille



Research team on interactions between

Light &



Matter

Organization of the talk



Quantum sensing with cold atoms

Experimental setup & observation of bistability

Towards driven-dissipative spin-squeezing & search for non-invasive witnesses

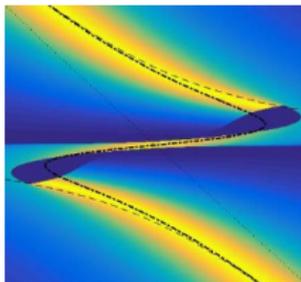
Optically dense clouds in cavities

Gravimetry with continuous monitoring of Bloch oscillations

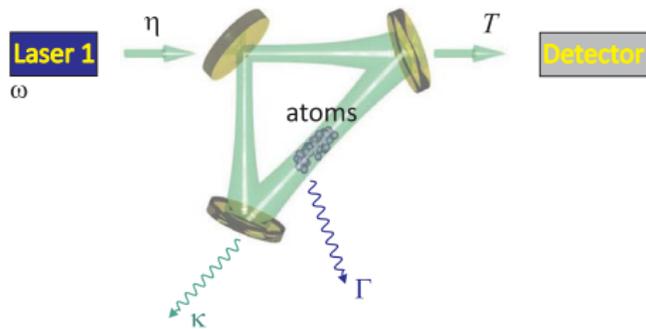


Saturation-induced resonant bistability in cold Strontium

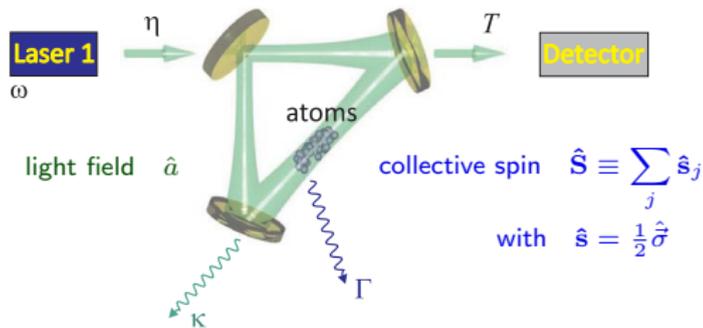
via non-linear interaction with a bad cavity



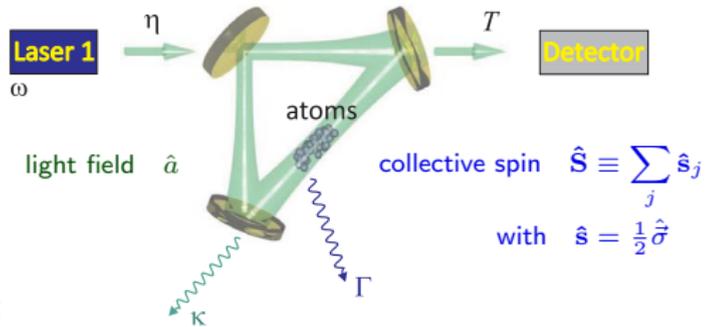
Dicke model



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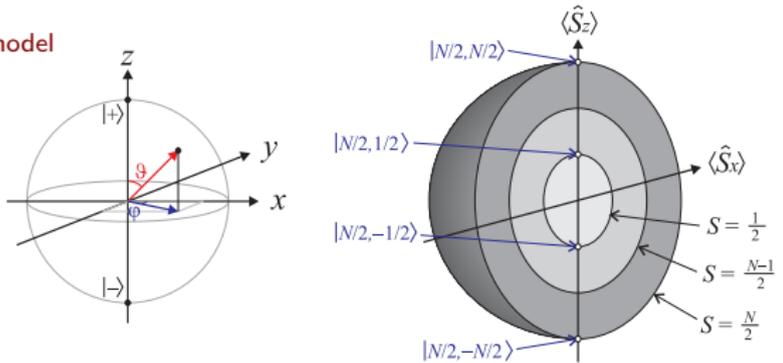
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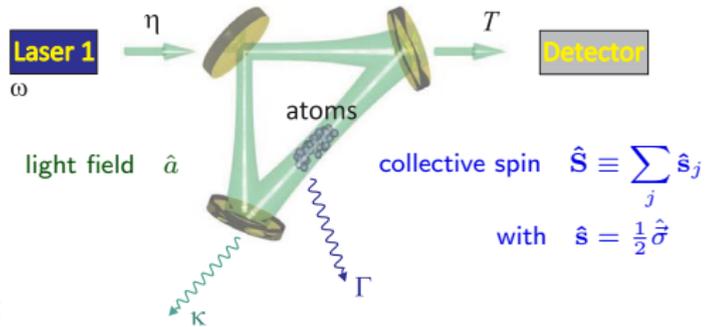
Atoms treated as non-interacting spins

no near field terms, only radiative coupling

coupled spin description \Rightarrow Dicke model



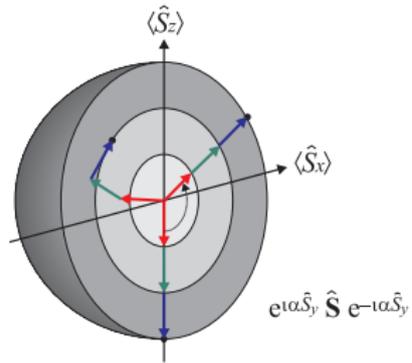
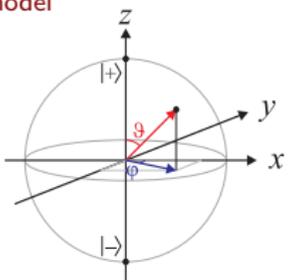
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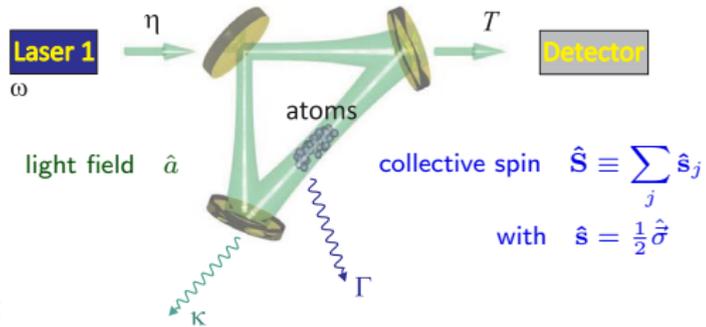
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Terms linear in $\hat{S}_{x,y,z}$ only perform rotations: $e^{i\alpha \hat{S}_z} \hat{S} e^{-i\alpha \hat{S}_z}$

\Rightarrow a coherent spin state always remains a coherent spin state

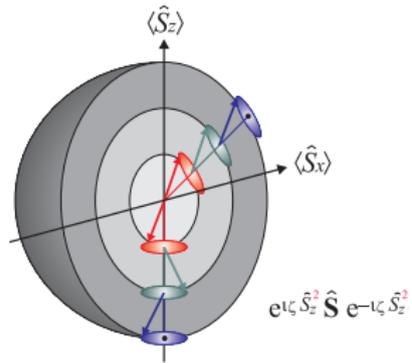
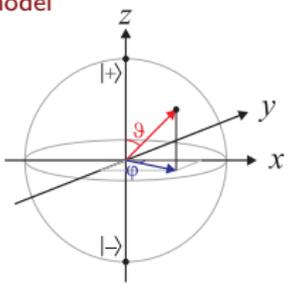
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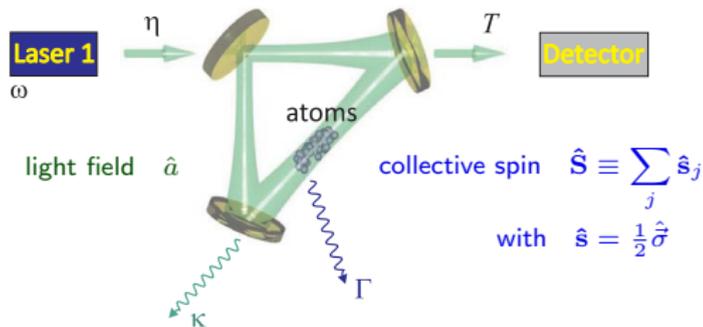
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\Rightarrow a coherent spin state always remains a coherent spin state

\Rightarrow no entanglement can be generated by linear spin operators in the Hamiltonian

Spin-squeezing requires non-linear terms: $e^{i\zeta \hat{S}_z^2} \hat{S} e^{-i\zeta \hat{S}_z^2}$

Why bad cavities?



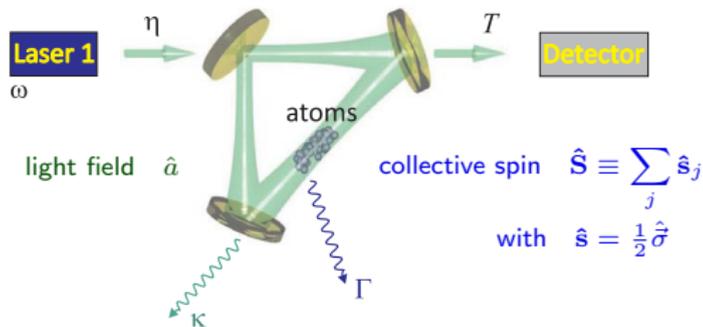
[Norcia, Lewis-Swan, Cline, Bihui Zhu, Rey, Thompson, Science **361**, 259 (2018)]

[Salvi, Poli, Vuletić, Tino, PRL **120**, 033601 (2018)]

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Why bad cavities?



Dicke model Hamiltonian $\hat{H} = -\Delta_c \hat{a}^\dagger \hat{a} - \eta \eta (\hat{a} - \hat{a}^\dagger) - \Delta_a (\frac{N}{2} + \hat{S}_z) + g(\hat{S}_+ \hat{a} + \hat{a}^\dagger \hat{S}_-)$

jump operator for dissipative processes $\hat{\mathcal{J}} = \kappa \hat{a}$ neglect spontaneous emission $\hat{\mathcal{J}} = \Gamma \hat{s}_j^- \rightarrow 0$

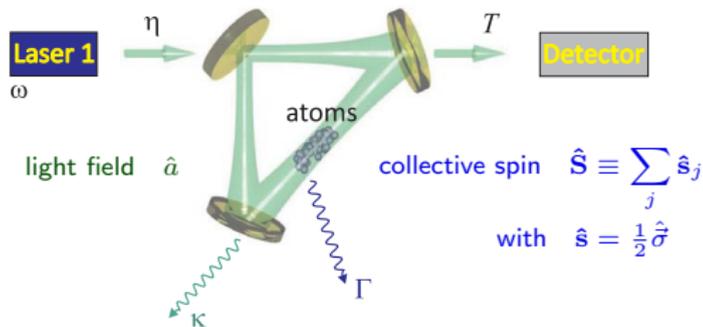
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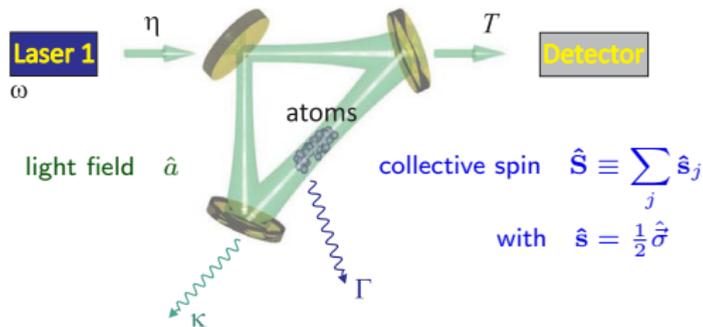
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bad-cavity approximation $\hat{H} \simeq -\frac{2}{g} \Im(\eta U_\kappa) \hat{S}_x - \frac{2}{g} \Re(\eta U_\kappa) \hat{S}_y - \Delta_a \hat{S}_z + U_c \hat{S}_+ \hat{S}_-$

collective jump operator $\hat{\mathcal{J}} = \kappa_c \hat{S}_-$

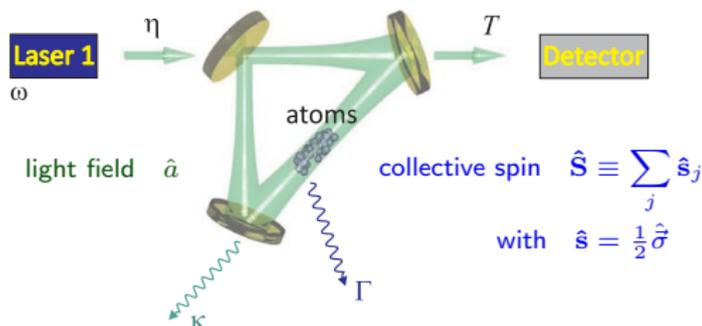
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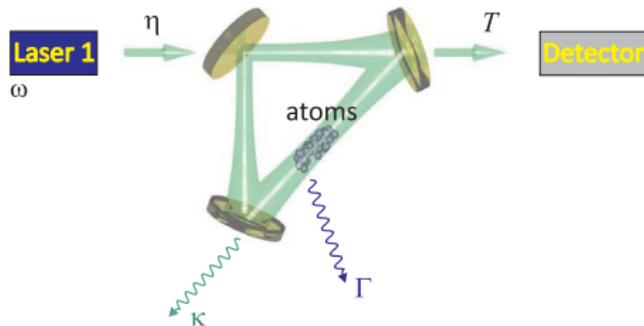
Lindbladian $\hat{\mathcal{L}}\hat{\rho} = \kappa_c (2\hat{S}_- \hat{\rho} \hat{S}_+ - \hat{\rho} \hat{S}_+ \hat{S}_- - \hat{S}_+ \hat{S}_- \hat{\rho})$
 \uparrow
 $\simeq \hat{S}_z^2$



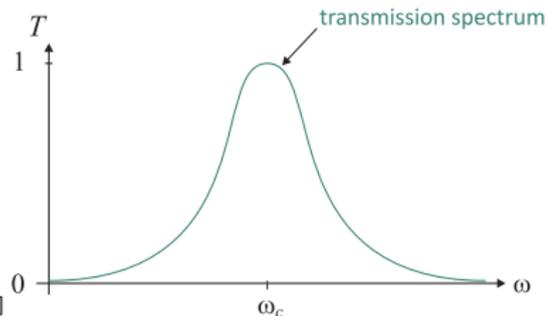
\implies non-linearity induced by feedback can generate entanglement

\implies spin squeezing and superradiant lasing

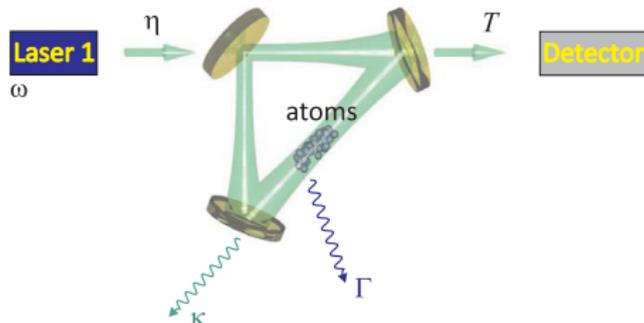
Storyboard for an experiment



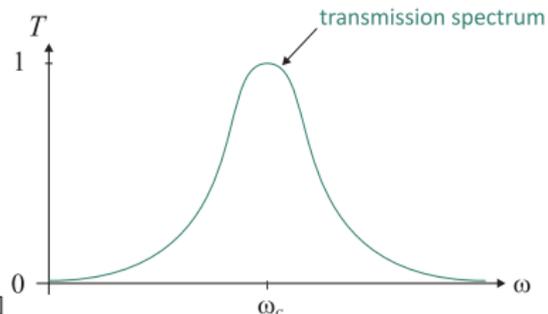
1) set up an experiment in the 'bad' cavity parameter regime ($\kappa \rightarrow \infty$)



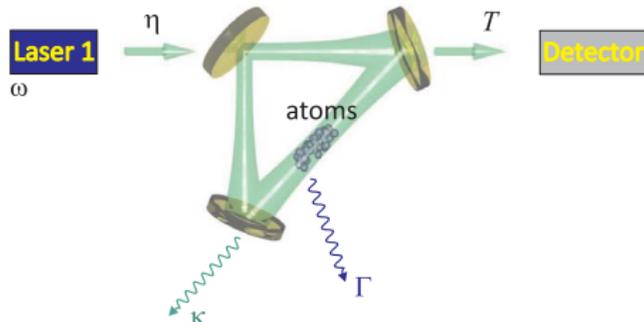
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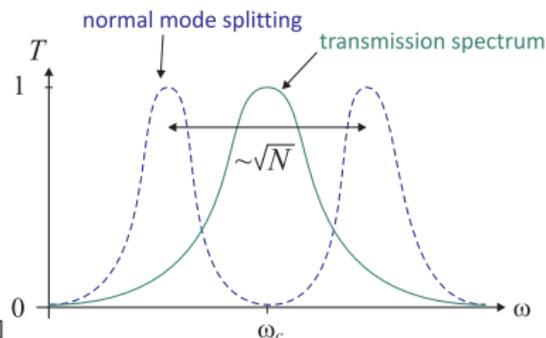
- 1) set up an experiment in the 'bad' cavity parameter regime ($\kappa \rightarrow \infty$)
- 2) take atoms with narrow transitions ($\Gamma \rightarrow 0$) and cool them



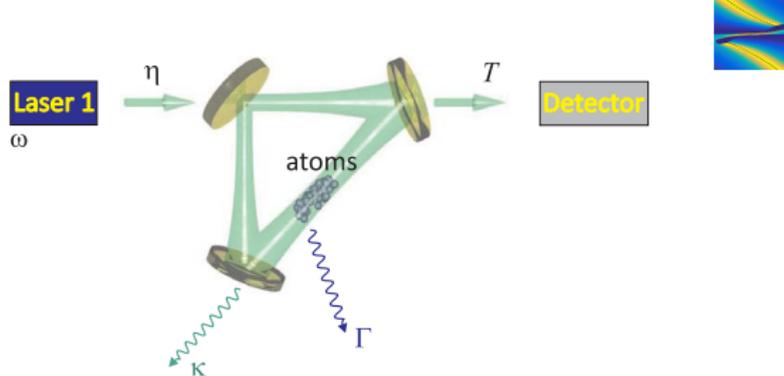
Storyboard for an experiment



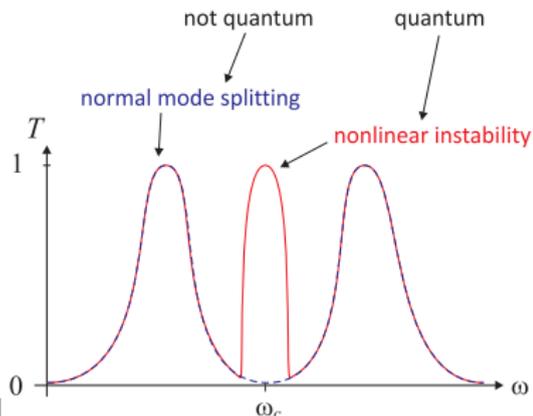
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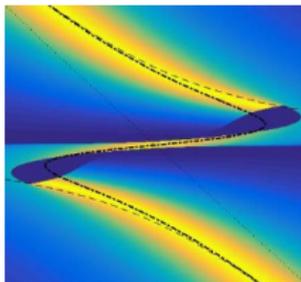


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- 4) verify non-linearity "on-resonance" ($\Delta_c = 0$)

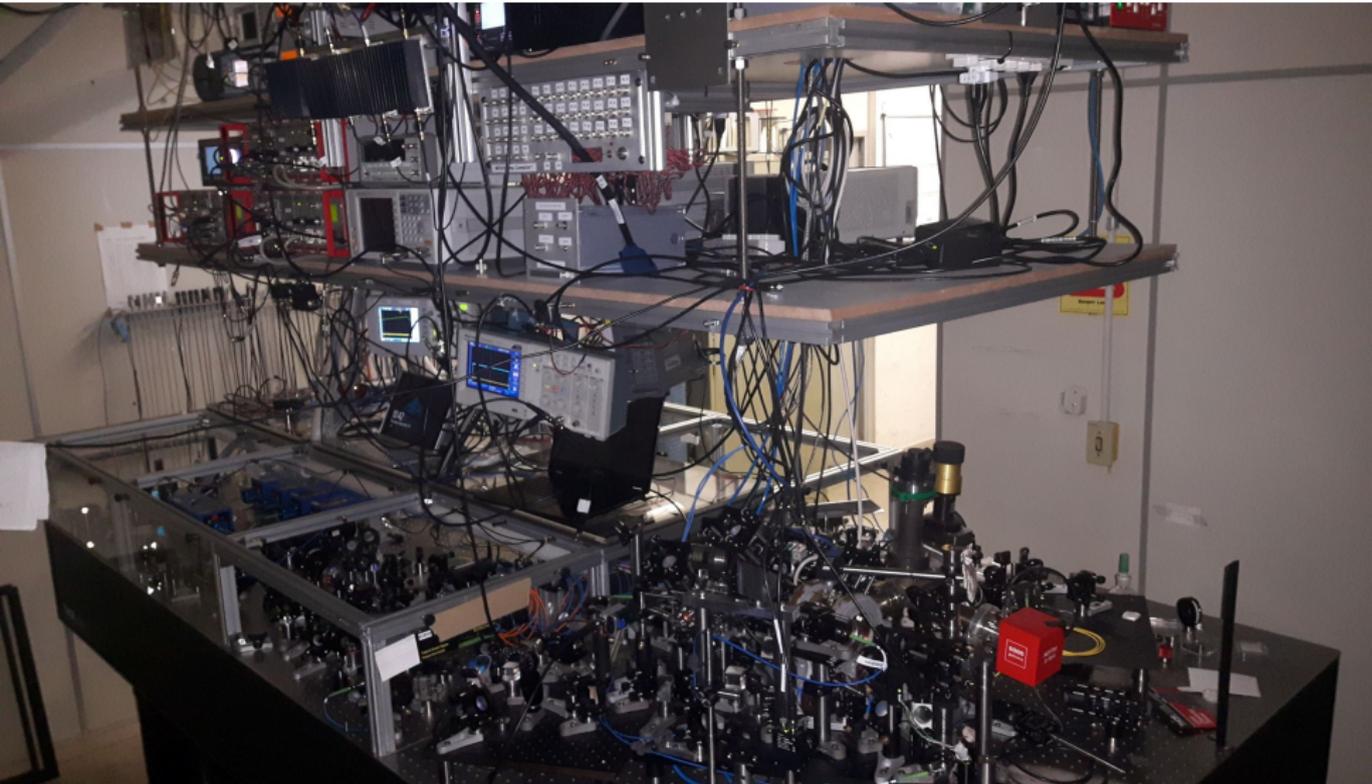


Saturation-induced resonant bistability in cold Strontium

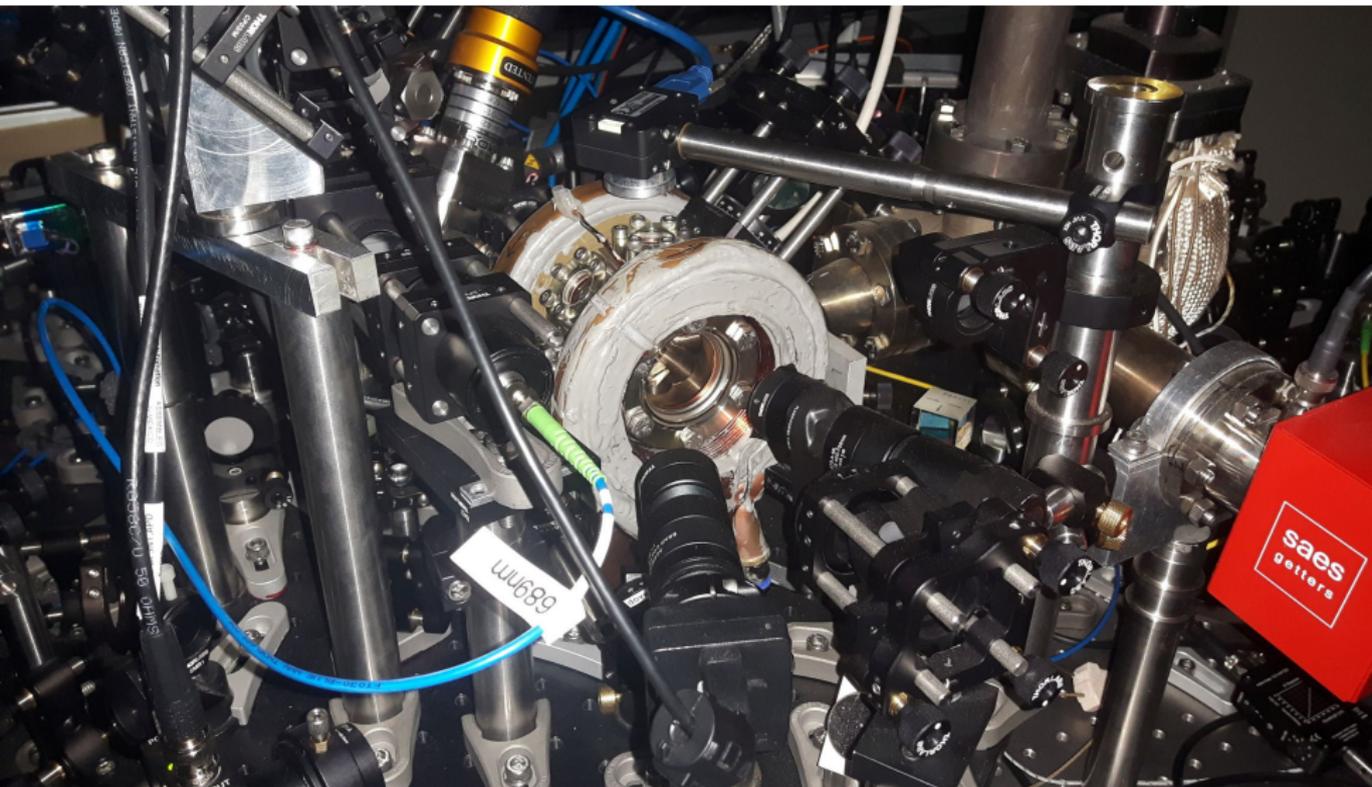
via non-linear interaction with a bad cavity



The experiment



The experiment



The experiment

strontium $\Gamma = 7.5 \text{ kHz}$

cavity decay $\kappa = 4.3 \text{ MHz}$



Experimental procedure & state of the art

experimental control

Strontium Control © Philippe W. Courteille

File Help

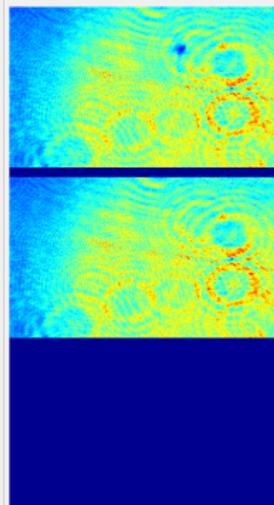


Switchboard



Absorption imaging

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Frames to queue: 3
Camera binning: 1 x 1
Camera origin: 0 x 0 pxl
Region of interest: 640 x 480 pxl
Number of pixels: 0 pxl



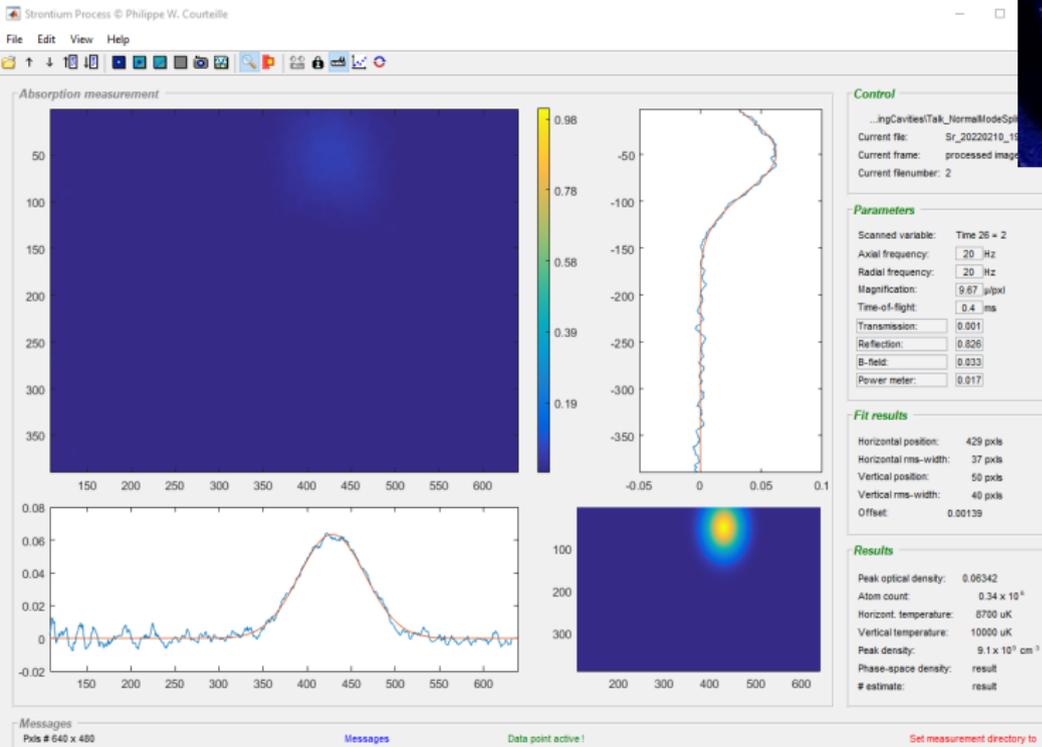
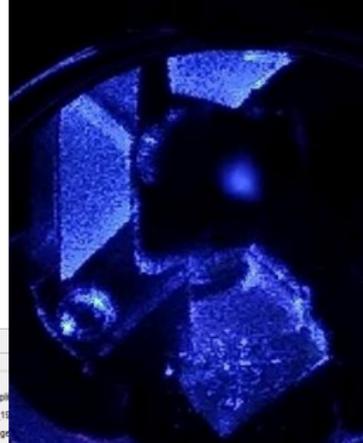
Messages

run / stop Points # 0 Loop # 0 Load settings from Sr_20220621_172638.par

No synthesizer available!

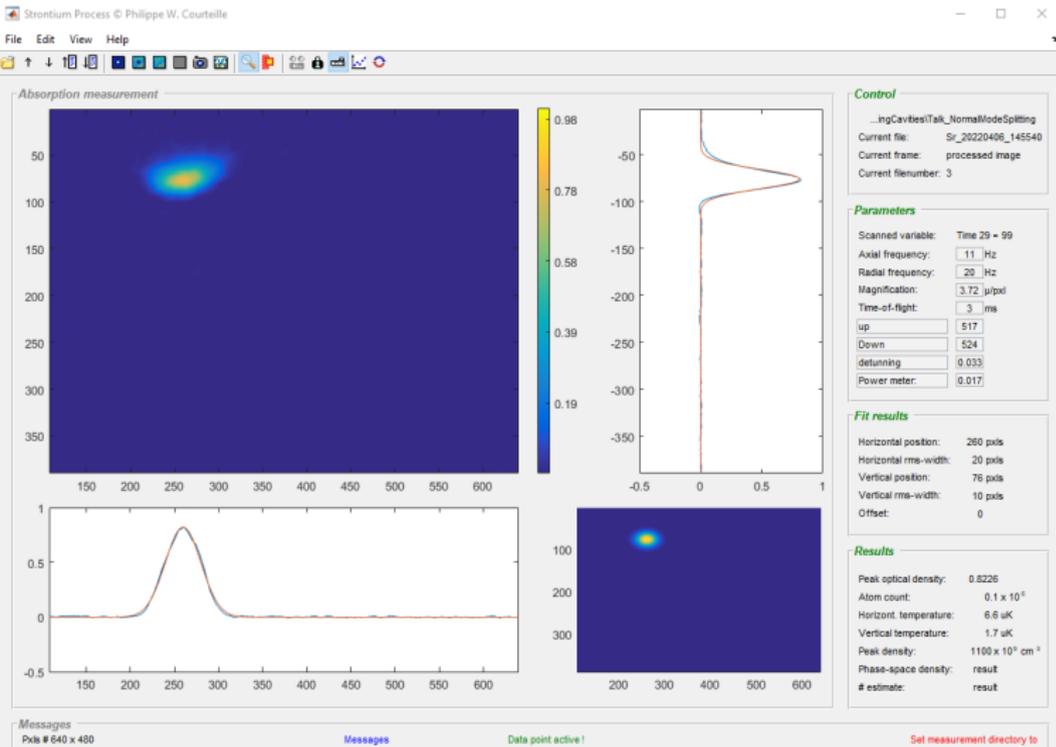
Experimental procedure & state of the art

trapping atoms in the blue MOT: $N = 10^6$ $T = 5$ mK



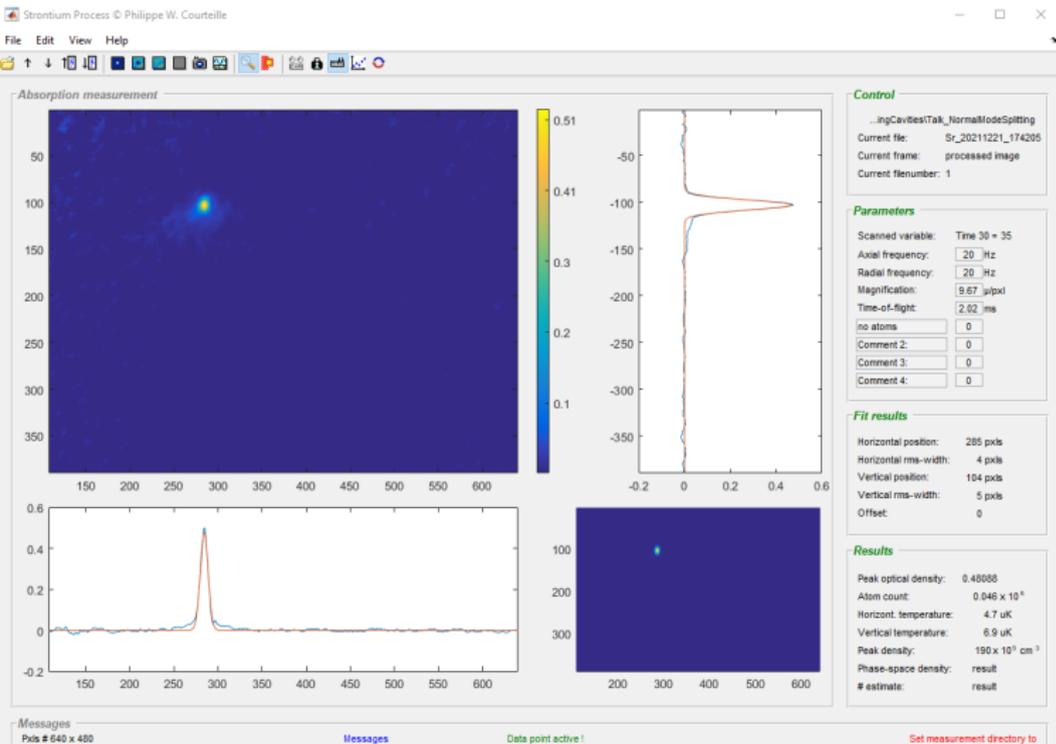
Experimental procedure & state of the art

cooling atoms in the red MOT



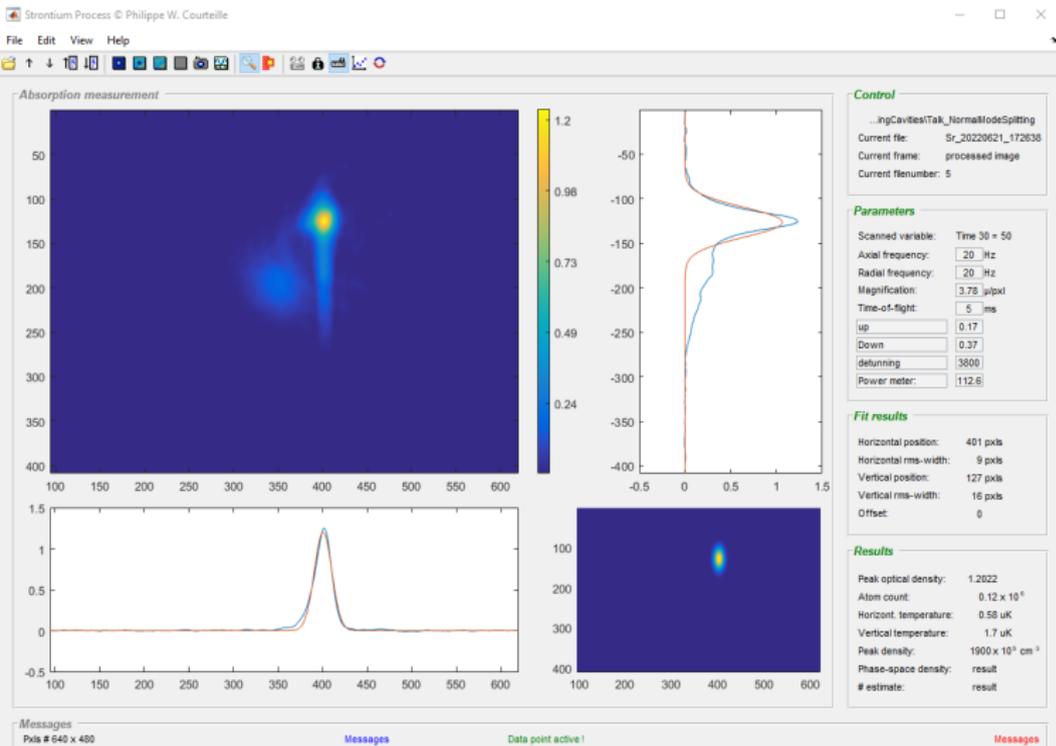
Experimental procedure & state of the art

cooling atoms in the red MOT: $N = 2 \cdot 10^5$ $T = 1 \mu\text{K}$



Experimental procedure & state of the art

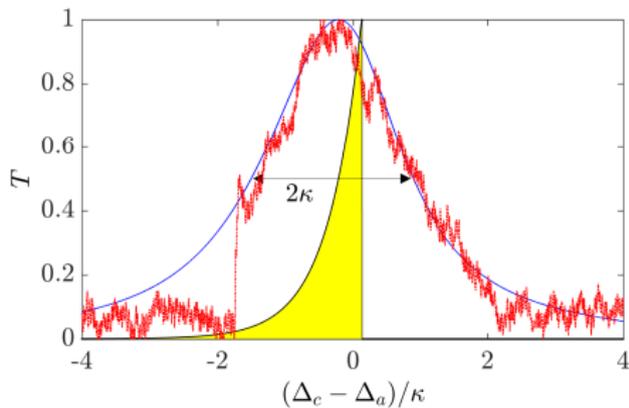
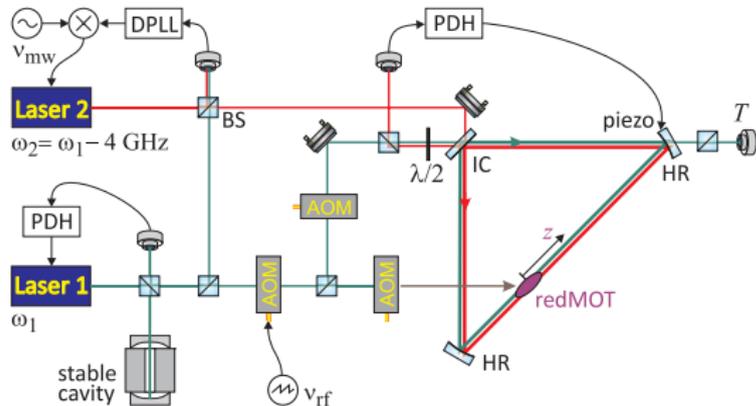
transferring atoms to the ring cavity mode via magnetic field ramp



Normal mode splitting



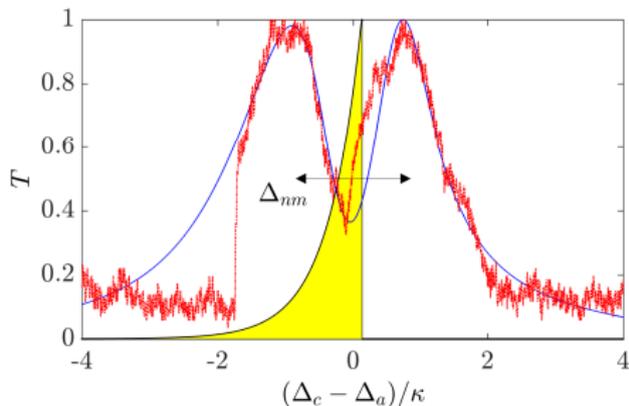
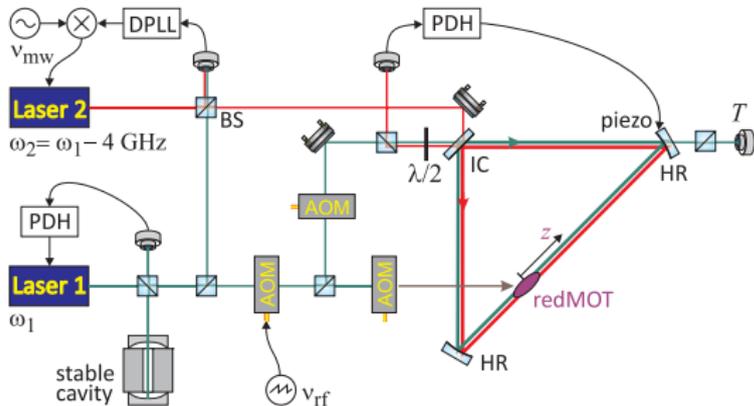
scanning laser frequency which pumps the cavity



Normal mode splitting

scanning laser frequency which pumps the cavity

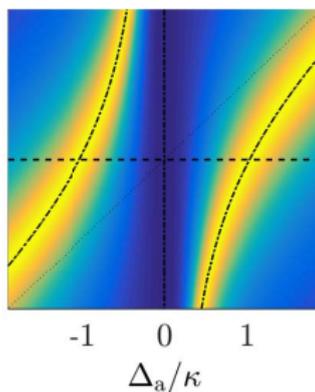
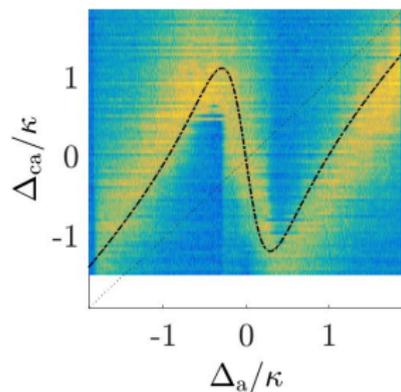
$$\Gamma \ll \kappa \lesssim g\sqrt{N} \equiv \Delta_{nm}$$



Normal mode splitting \equiv 1D photonic band gap



avoided crossing + instable feature



$$\Delta_{ca} \equiv \Delta_a - \Delta_c$$

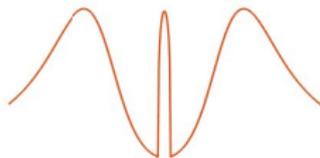
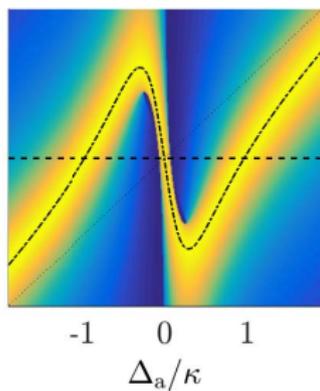
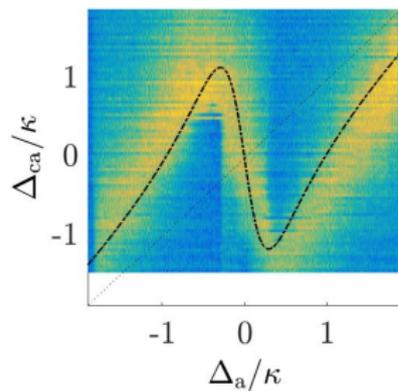
$$\Delta_c = \frac{Ng^2 \Delta_a}{\Delta_a^2 + \Gamma^2/4}$$



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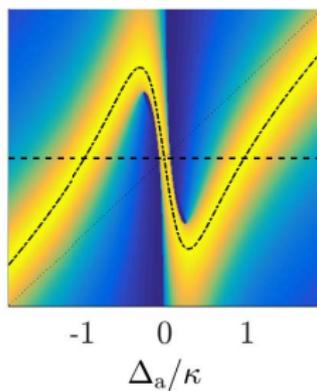
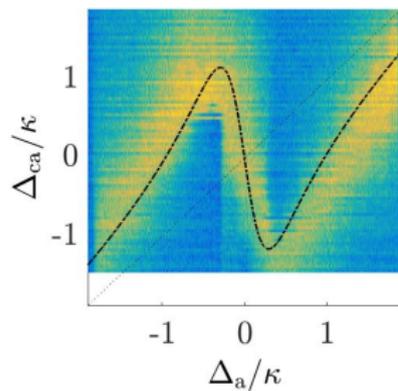
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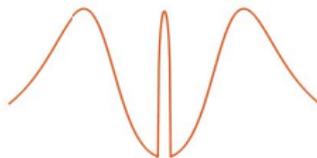


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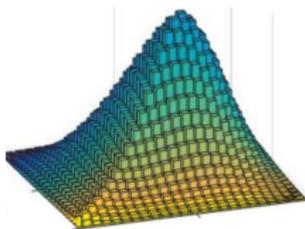


adiabatic elimination only near $\Delta_a = 0$



Route to steady state spin-squeezing

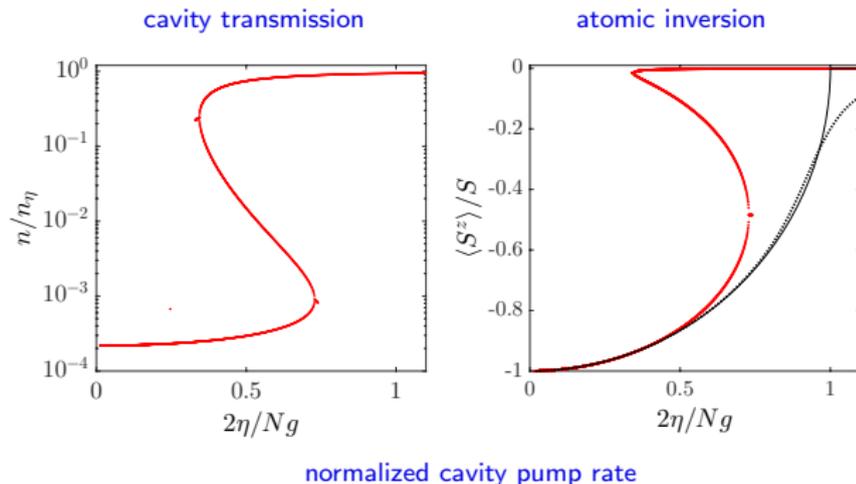
and search for spin-squeezing witnesses



Steady state behavior within mean field



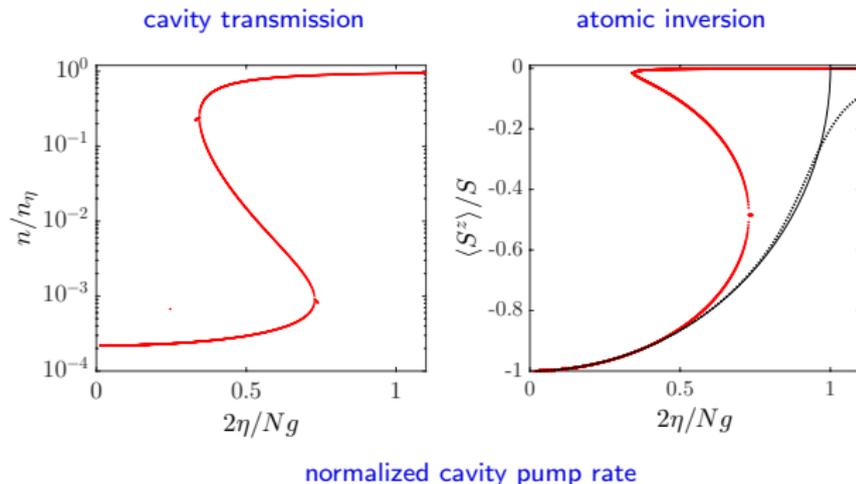
bistability curve for $\Delta_a = 0 = \Delta_c$ and $\langle \hat{S}_\pm \hat{S}_z \rangle = \langle \hat{S}_\pm \rangle \langle \hat{S}_z \rangle$ and $\frac{d}{dt} \hat{a} = 0 = \frac{d}{dt} \hat{S}$



Steady state behavior within mean field



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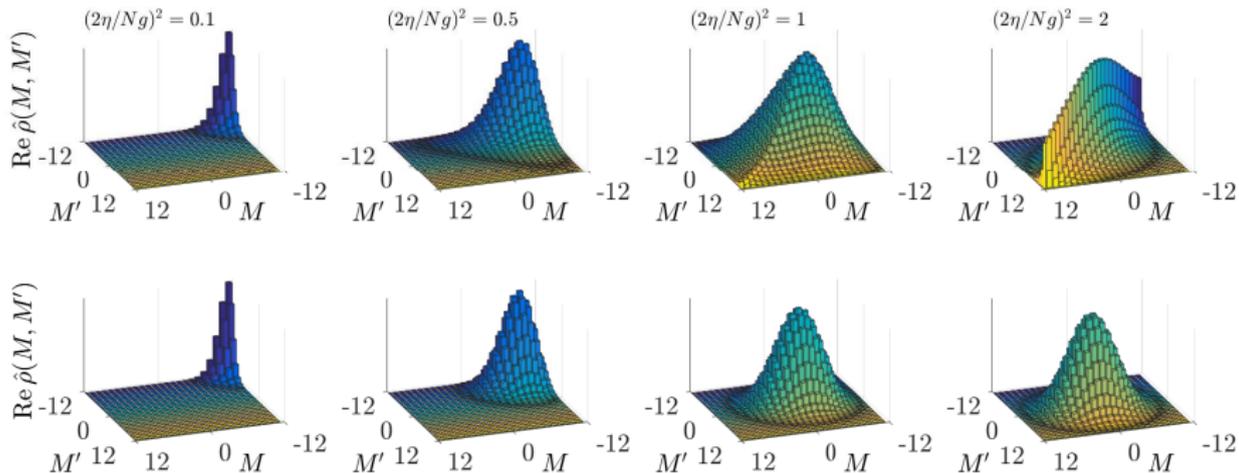


cooperativity $\Upsilon_N = \frac{Ng^2}{\kappa\Gamma} = \frac{N\kappa_c}{\Gamma} = \frac{\text{collective decay}}{\text{single-atom decay}}$

Beyond mean field without spontaneous emission



driven-dissipative steady state density matrix in Dicke basis

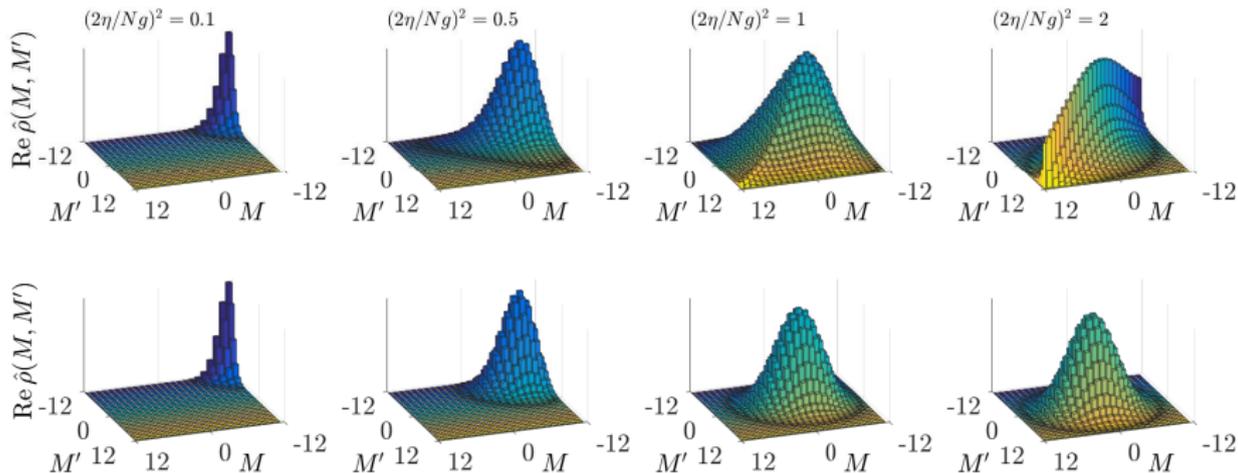


corresponding coherent spin state

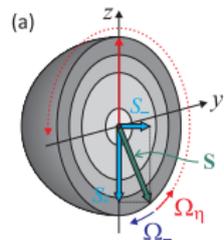
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corresponding coherent spin state

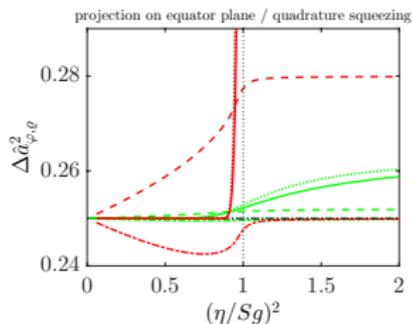
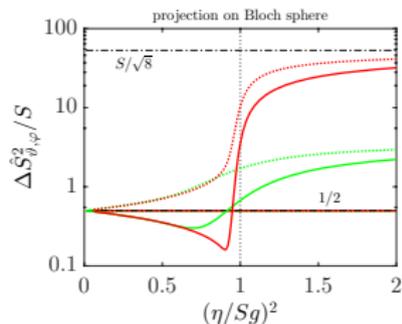
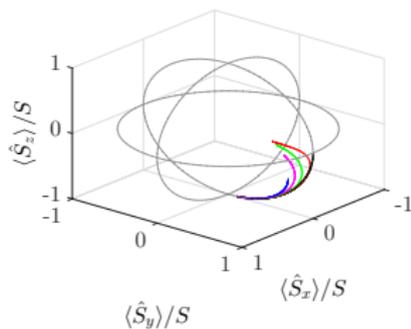
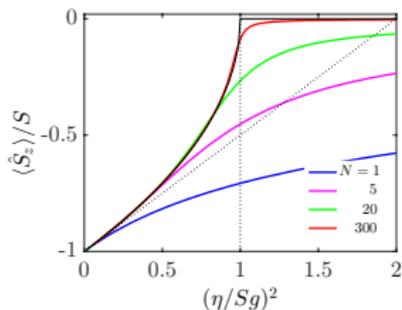


Dicke phase transition



non-linearity provided by **collective dissipation + pumping** rather than Hamiltonian evolution

$$\dot{\hat{\rho}} = \imath[\hat{\rho}, \hat{H}] + \mathcal{L}\hat{\rho} \quad \text{with} \quad \hat{H} = \frac{\imath\eta g}{\kappa} \hat{S}_x \quad \text{and} \quad \mathcal{L}\hat{\rho} = \frac{g^2}{\kappa} (2\hat{S}_- \hat{\rho} \hat{S}_+ - \hat{S}_+ \hat{S}_- \hat{\rho} - \hat{\rho} \hat{S}_+ \hat{S}_-)$$

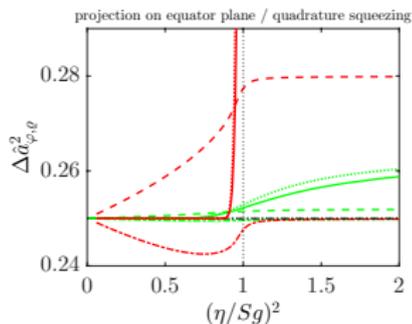
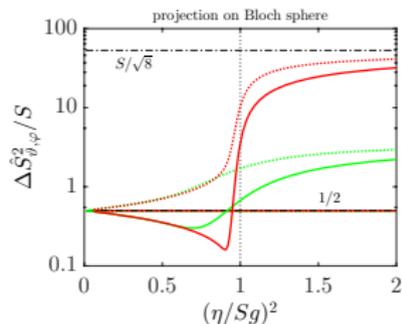
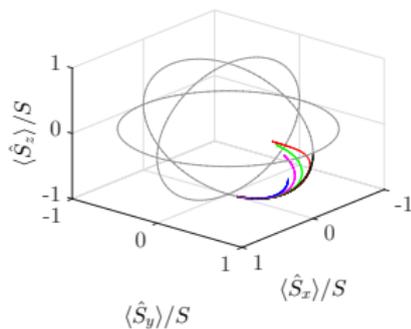
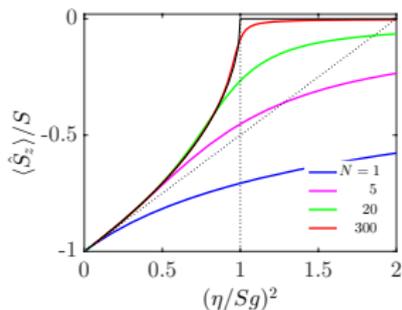


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driven-dissipative spin-squeezing

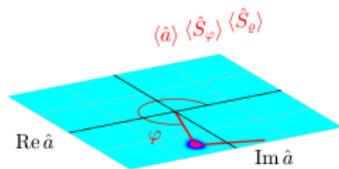
light squeezing

Coherently radiating spin-squeezed states

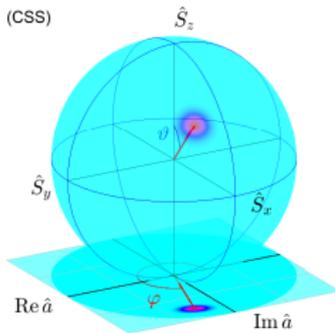
$$\hat{a}_{\varphi,\varrho} = \hat{c} - \imath \frac{g}{\kappa} \hat{S}_{\varrho,\varphi}$$

$$\Delta \hat{a}_{\varphi,\varrho}^2 - \frac{1}{4} = \frac{g^2}{\kappa^2} (\Delta \hat{S}_{\varrho,\varphi}^2 + \frac{1}{2} \langle \hat{S}_z \rangle)$$

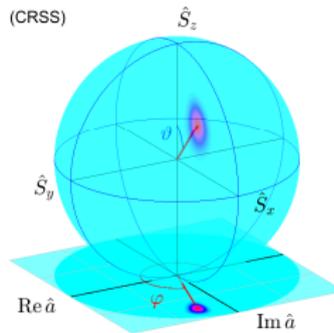
(CLS)



(CSS)



(CRSS)



[Wang, Wu et al., New J. Phys. **16**, 063039 (2014)]

[Somech, Leppenen, Shahmoon, et al., PRA **108**, 0203725 (2023) & PRX Quantum **5**, 010349 (2024) & arXiv:2404.02134]

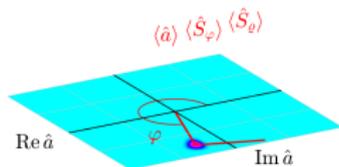
[Song, Rey, Thompson et al., Science Adv. **11**, eadu5799 (2025)]

Coherently radiating spin-squeezed states

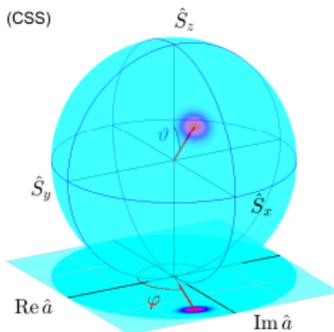
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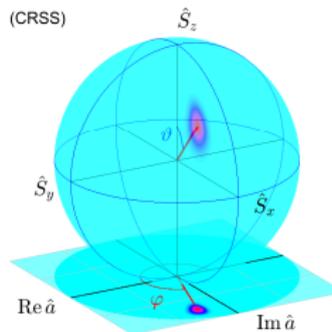
(CLS)



(CSS)



(CRSS)



optical spin-squeezing witness?

[Wang, Wu et al., New J. Phys. **16**, 063039 (2014)]

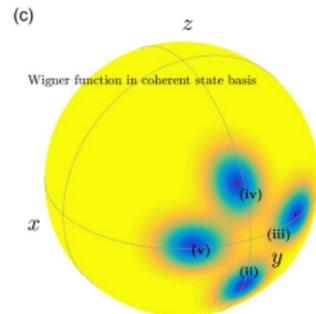
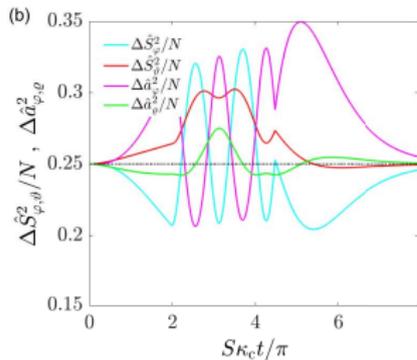
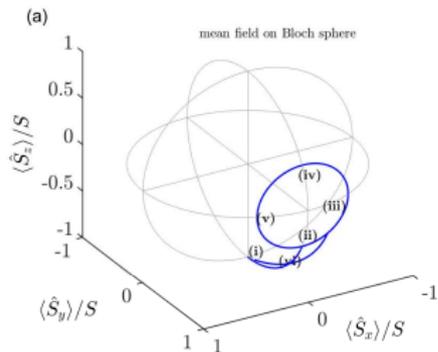
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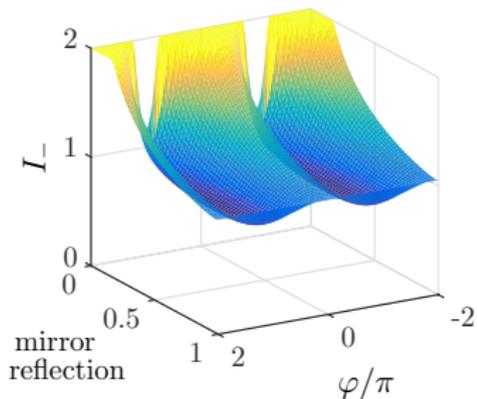
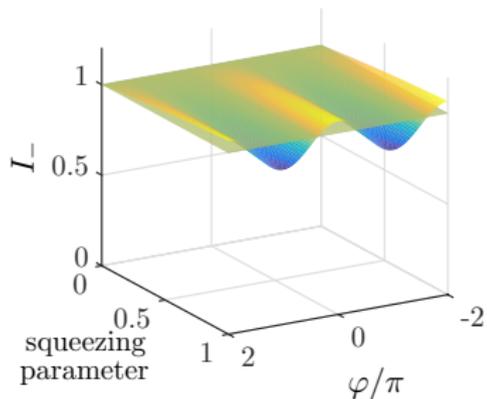
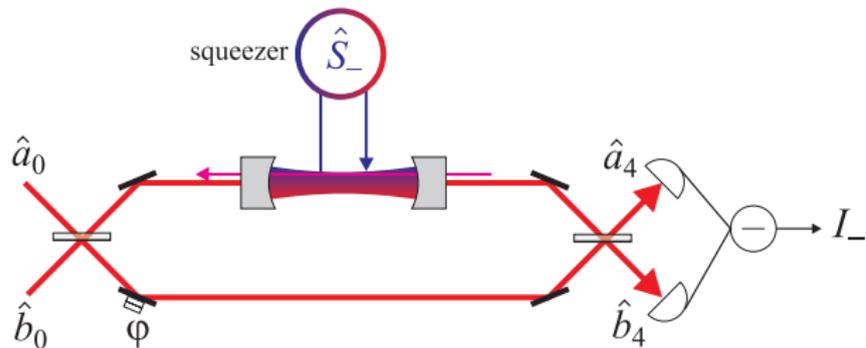
Optical spin-squeezing witness



rotation pulse for squeezing axis works, but only for times short compared to $\kappa_c = \frac{g^2}{\kappa} \approx (2\pi) 20 \text{ Hz}$



Homodyne detection



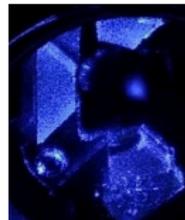
contrast reduction due to losses

Quintessence for bistability



Done:

- bistability observed on resonance with a 'bad cavity'! \implies non-linearity



[Meiser et al., PRL **102**, 163601 (2009)]

[Debnath, Zhang, Mølmer, PRA **98**, 063837 (2018)]

[Rosario, Santos, Piovella, Kaiser, Cidrim, R. Bachelard, PRL **133**, 050203 (2024)]

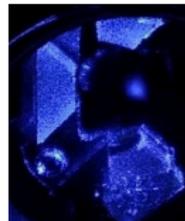
[recent work of groups of Vuletic, Schleier-Smith, Thompson, Rey, ...]

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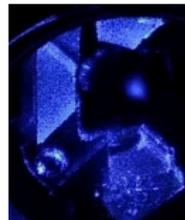
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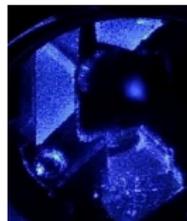
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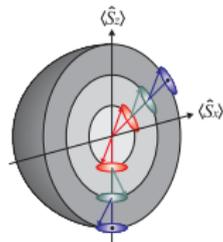
non-linearity + quantumness \implies new ideas on squeezing or superradiant lasing

To do:

drive atoms into spin-squeezed steady state

find optical spin-squeezing witnesses

generate inversion $> 50\%$ (e.g. via optical pumping) for light amplification



[Meiser et al., PRL **102**, 163601 (2009)]

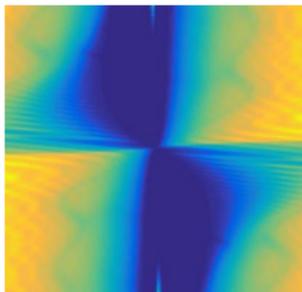
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[recent work of groups of Vuletic, Schleier-Smith, Thompson, Rey, ...]

Optically dense clouds in optical lattices

Cavity models breaking down

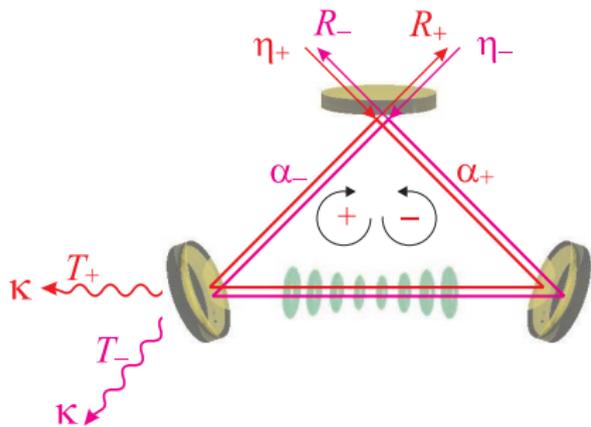


Optically dense clouds in optical lattices

atom numbers $N \approx 200000$

lattice sites $N_s \approx 300$

optical density $OD = \frac{6N}{k^2 w^2} \approx 3$



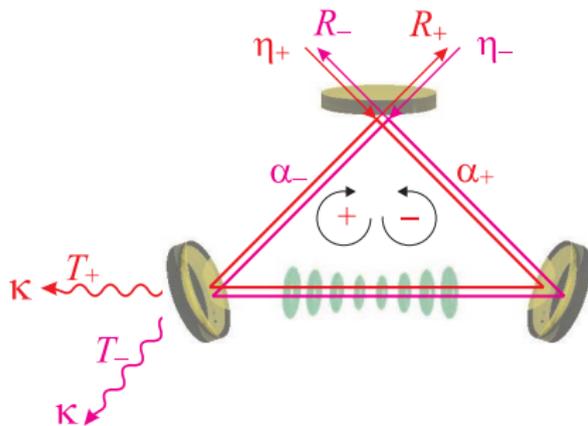
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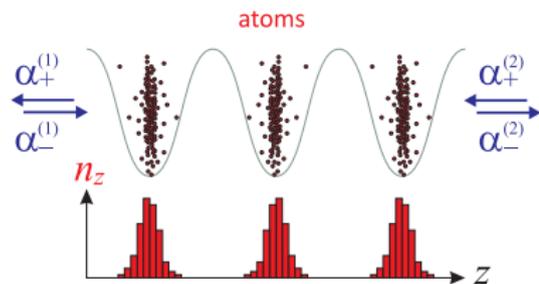


⇒ absorption and multiple scattering

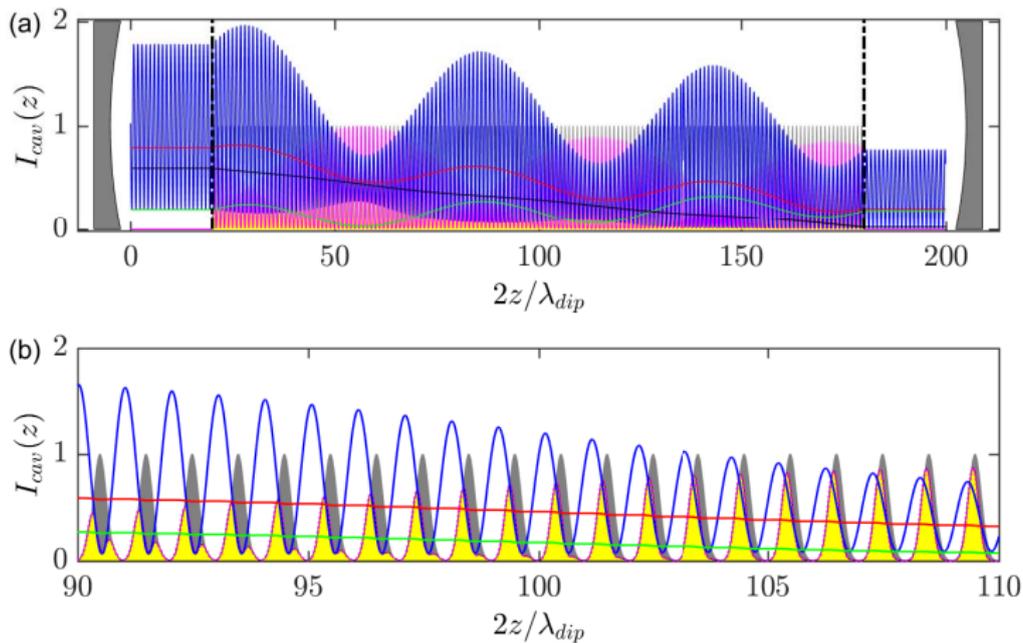
⇒ cavity concept invalid (no valid cavity mode function)

⇒ Open Dicke Model (ODM) no longer valid

⇒ use Transfer Matrix Model (TMM)



Intracavity intensity



[Deutsch, Spreuw, Rolston, Phillips, PRA **52**, 1394 (1995)]

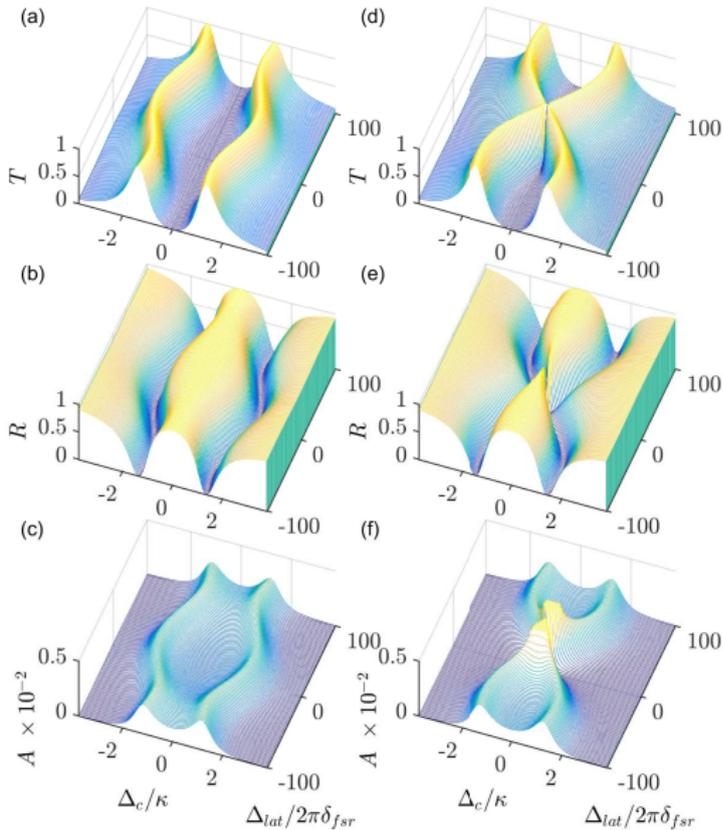
[Slama, von Cube, Kohler, Zimmermann, Courteille, Phys. Rev. A **73**, 023424 (2006)]

[Schilke, Zimmermann, Courteille, Guerin, Nature Phot. **6**, 101 letter (2012)]

[Samoylova, Piovella, R. Bachelard, Ph.W. Courteille, Opt. Comm. **312**, 94 (2014)]

[Rivero, de França, Pessoa, Teixeira, Slama, Courteille, New J. Phys. **25**, 093053 (2023)]

Mode mode splitting \equiv 1D photonic band gap

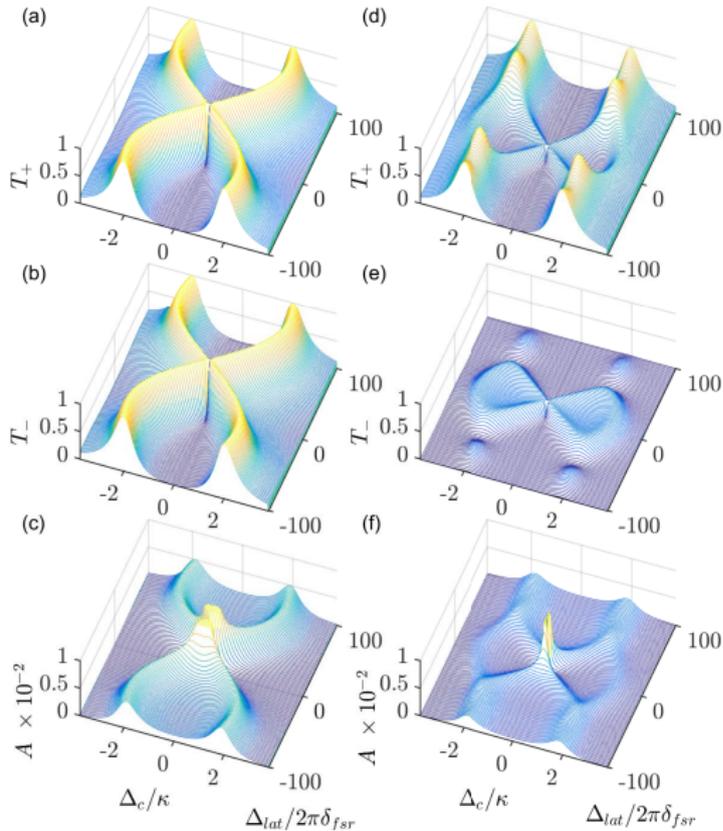


linear cavity

left/right: atoms sit at antinodes/nodes

$OD < 1 \Rightarrow$ ODM = TMM

Normal mode splitting \equiv 1D photonic band gap

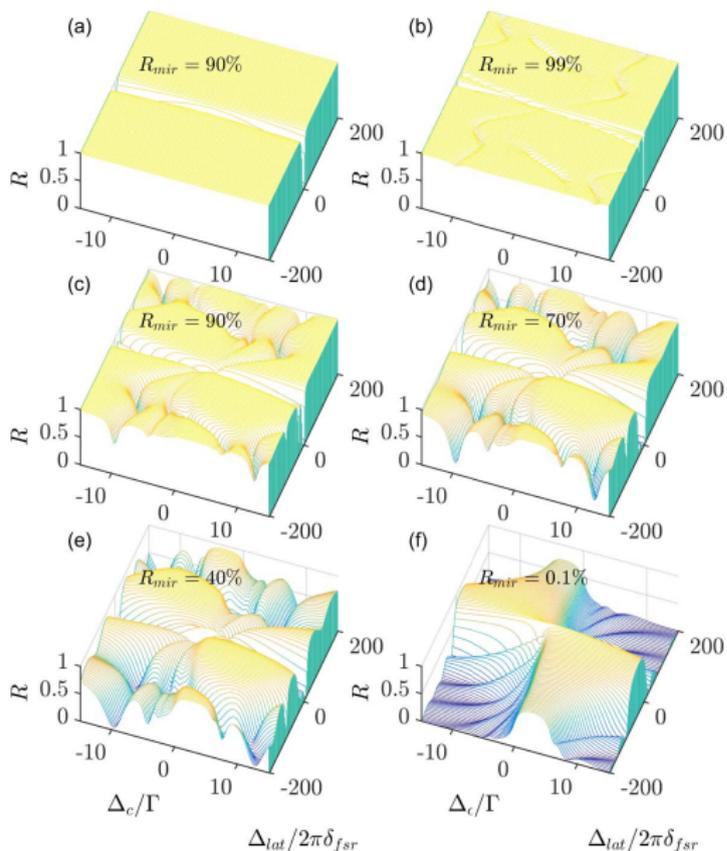


ring cavity

left/right: atoms sit at antinodes/nodes

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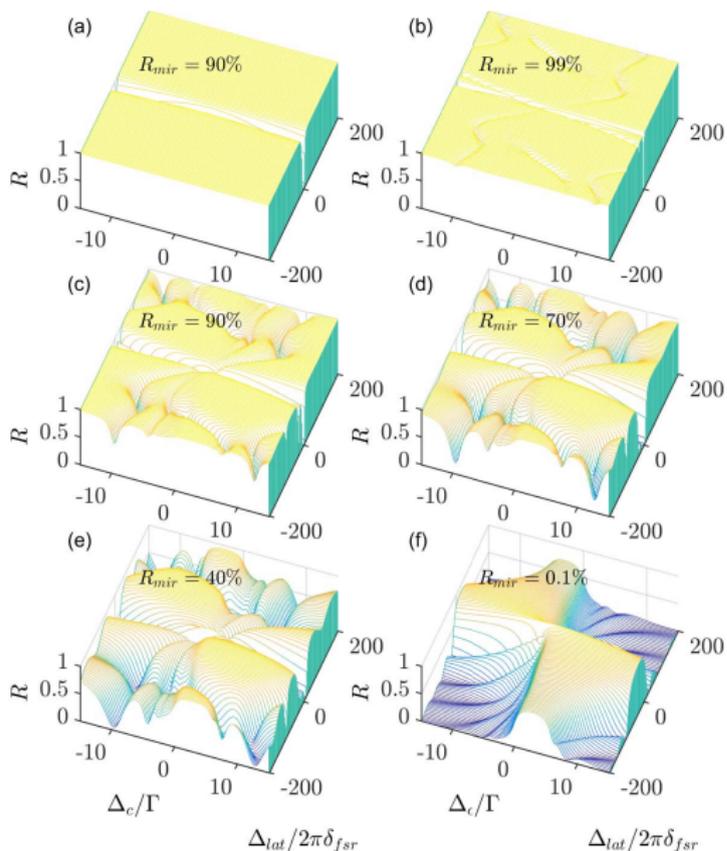


linear cavity

$OD > 1 \Rightarrow ODM \neq TMM$

\Rightarrow appearance of photonic bandgap

Normal mode splitting \equiv 1D photonic band gap



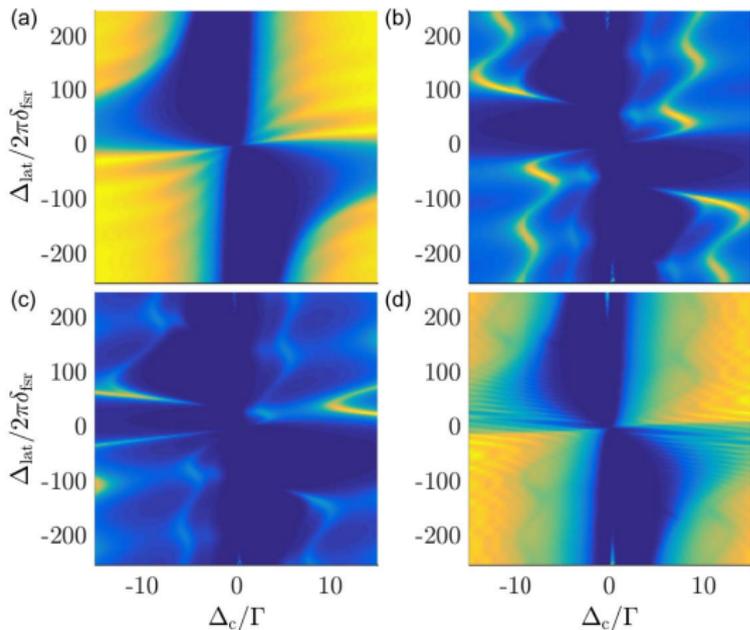
linear cavity

$OD > 1 \Rightarrow ODM \neq TMM$

\Rightarrow appearance of photonic bandgap

\Rightarrow notion of cavity mode function fails

Dense cloud with macroscopic boundary conditions



cavity = filter for specific reflection paths

Quintessence for dense clouds



need a full quantum model working at any ' I_{sat} ' and ' OD '

Quintessence for dense clouds



need a full quantum model working at any ' I_{sat} ' and ' OD '

ODM and TMM neglect direct photon exchange via radiation modes (\mathbf{k}, λ)

$$\hat{H}_{\text{Ising}} = \sum_{i \neq j} \Delta_{ij} \hat{\sigma}_j^+ \hat{\sigma}_i^-$$

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treat mirrors as macroscopic boundaries \implies cooperative environment shaping $g_{\mathbf{k}\lambda}$

$$\hat{H} = \hbar \sum_{\mathbf{k}, \lambda} \sum_j (\hat{\sigma}_j^+ e^{-ikz_j} + \hat{\sigma}_j^- e^{ikz_j}) (g_{\mathbf{k}\lambda} \hat{a}_{\mathbf{k}\lambda} + g_{\mathbf{k}\lambda}^* \hat{a}_{\mathbf{k}\lambda}^\dagger)$$

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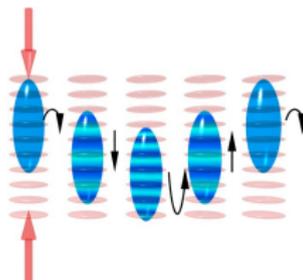
JCM with 1 atom breaks down for Purcell factor $\gg 1$

$$OD = \sigma_0 n_{\text{at}} N_s \frac{\lambda_{\text{lat}}}{2} = \frac{6N}{k^2 w^2} = \frac{\pi Y N}{F} > 1$$



Matter wave Bloch oscillations

for inertial sensing

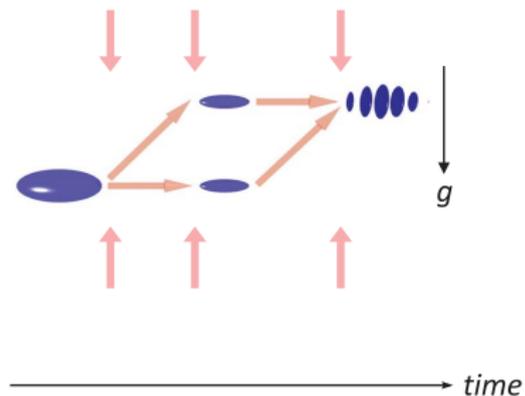


Gravimetry with Bose-Einstein condensates



differential phase shift of de Broglie waves

→ matter wave interferometers

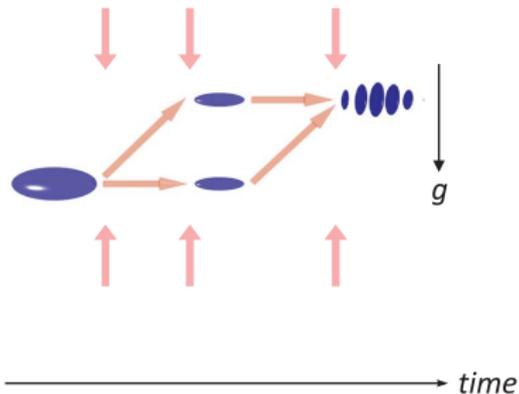




Gravimetry with Bose-Einstein condensates

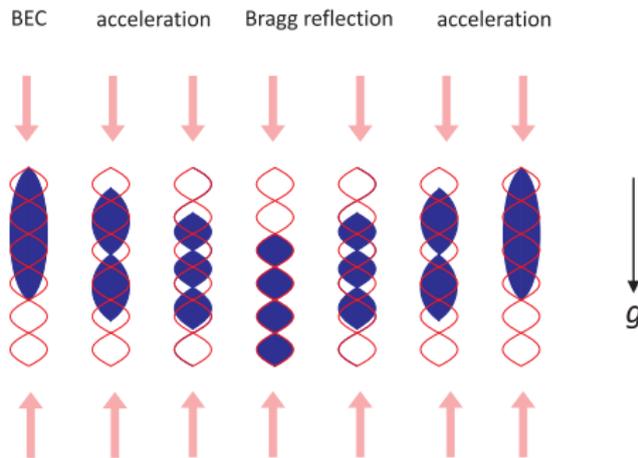
differential phase shift of de Broglie waves

→ matter wave interferometers



matter wave Bloch oscillations in a periodic potential

- wavelength $\lambda_{dB} = \frac{h}{mv}$

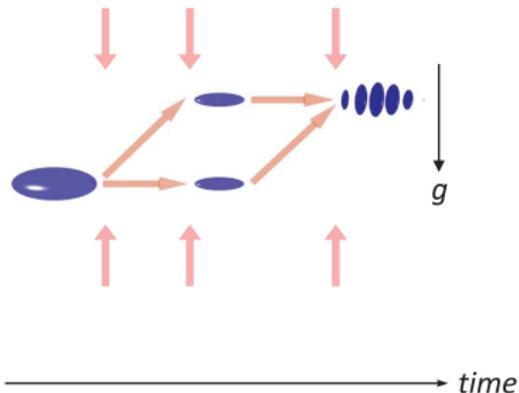




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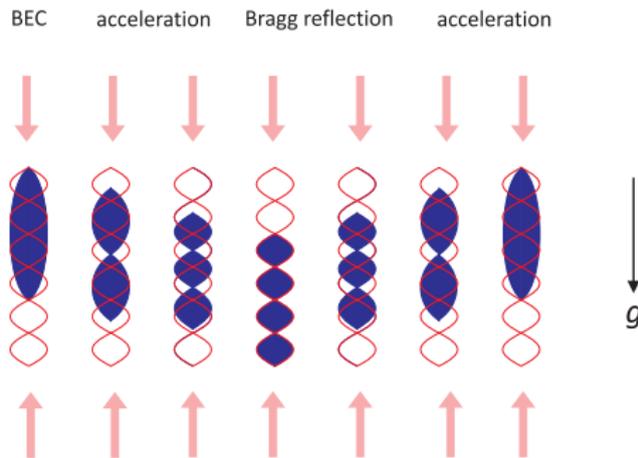


matter wave Bloch oscillations in a periodic potential

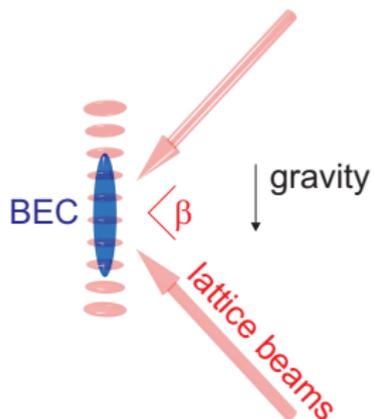
• wavelength $\lambda_{dB} = \frac{h}{mv}$

• frequency $\nu_{Bloch} = \frac{mg}{2\hbar k}$

→ measure gravity g



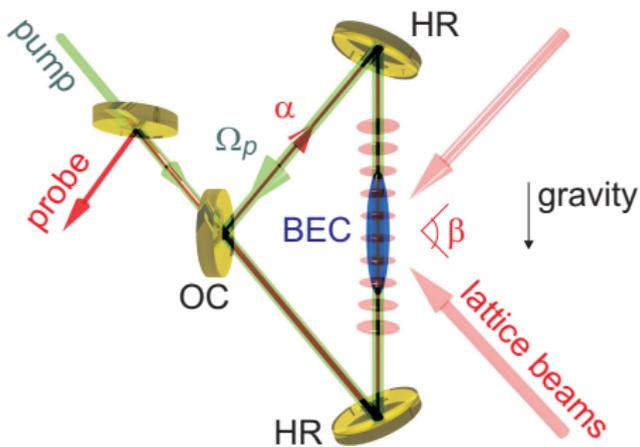
Continuous monitoring Bloch oscillations in a cavity



[Samoylova, Piovella, Robb, Bachelard, Courteille, Opt. Exp. **23**, 14823 (2015)]

[Samoylova, Piovella, Hunter, Robb, Bachelard, Courteille, Las. Phys. Lett. **11**, 126005 (2014)]

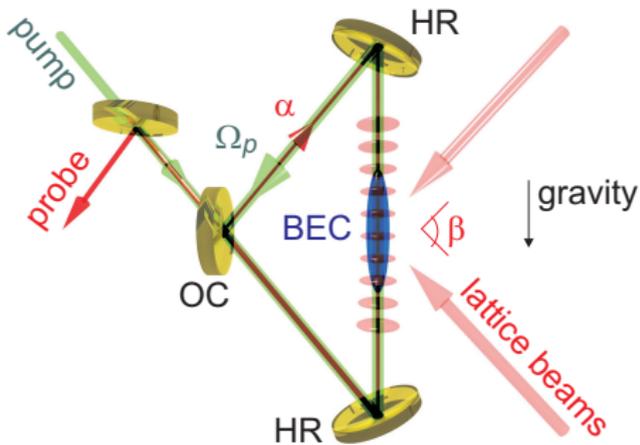
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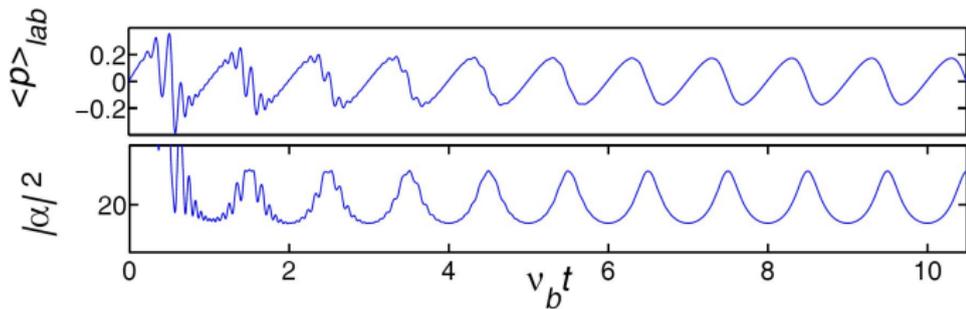


$$N = 8 \cdot 10^4$$

$$\kappa = 160\omega_{\text{rec}}$$

$$U_0 = 0.04\omega_{\text{rec}}$$

$$W_0 = 80U_0$$



[Samoylova, Piovella, Robb, Bachelard, Courteille, Opt. Exp. **23**, 14823 (2015)]

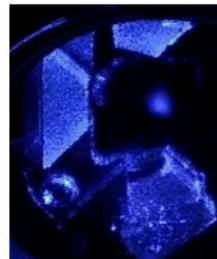
[Samoylova, Piovella, Hunter, Robb, Bachelard, Courteille, Las. Phys. Lett. **11**, 126005 (2014)]

Quintessence



Done:

- 200000 atoms cooled down to single-photon recoil limit

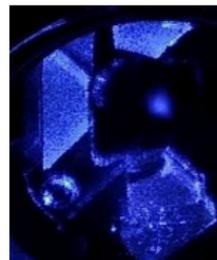


Quintessence



Done:

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- confinement in an optical lattice sustaining 1 trapped state

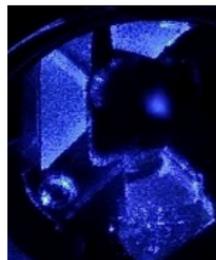


Quintessence



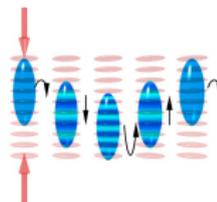
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To do:

search for signatures of Bloch oscillations in light modes

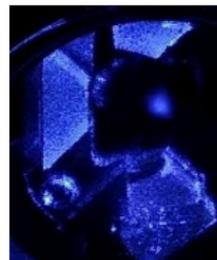


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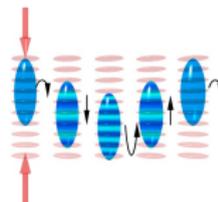
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To do:

search for signatures of Bloch oscillations in light modes

continuous monitoring of gravity



The team

Raul Teixeira, Dalila Rivero, Gustavo de França, Claudio Pessoa, Thiago da Silva

Ana Cipris, Matheus Rodrigues, Daniel Coelho, Felipe Brambila, Thales Pereira



Linearity of quantum mechanics

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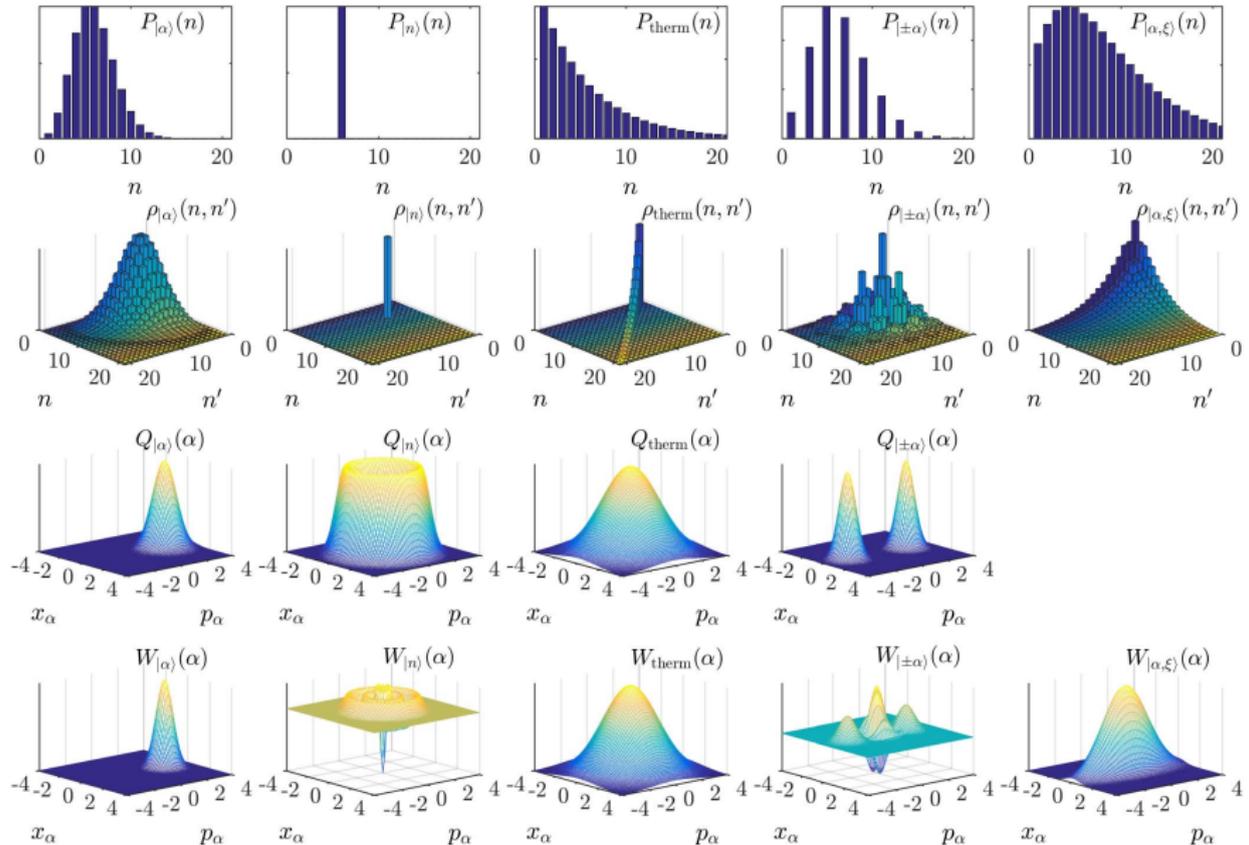
QM is a Lie algebra \implies linearity \implies our universe is a quantum computer

$V(\mathbf{r})$ is an artifact [just like $n_{\text{refr}}(\mathbf{r})$ in optics or $\Phi(\mathbf{r})$ and $\mathbf{A}(\mathbf{r})$ in E-dynamics]

need to break down physical processes into momentum-conserving collisions \implies microscopic description of V

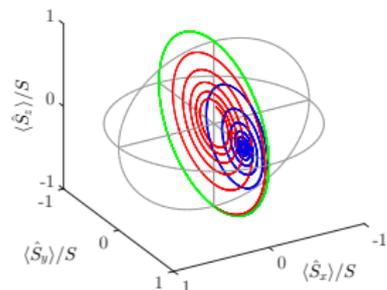
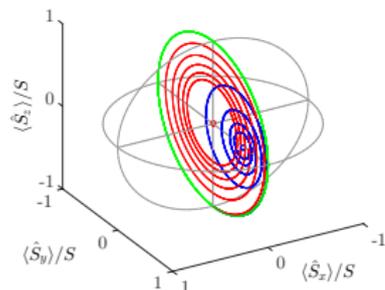
Space is homogeneous \implies momentum always conserved

Representation of particular states of light



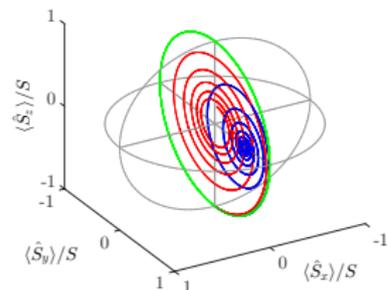
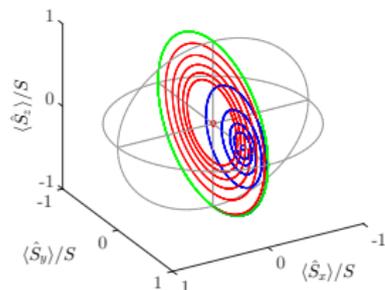
Impact of spontaneous emission

spontaneous emission recovers steady state excitation

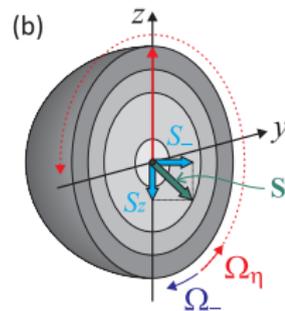
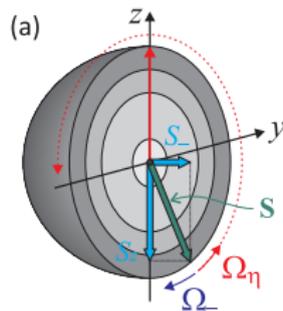


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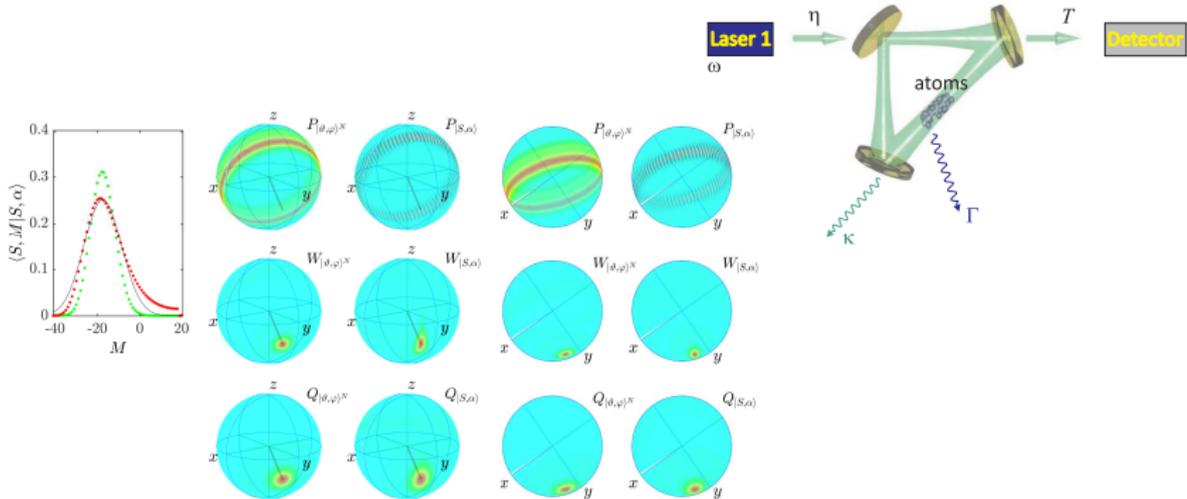


driven-damped rigid rotor flipping over



Beyond mean field: Coherently radiating spin-squeezed state

non-linearity provided by **collective dissipation + pumping** rather than Hamiltonian evolution



[Wang, Wu et al., New J. Phys. **16**, 063039 (2014)]

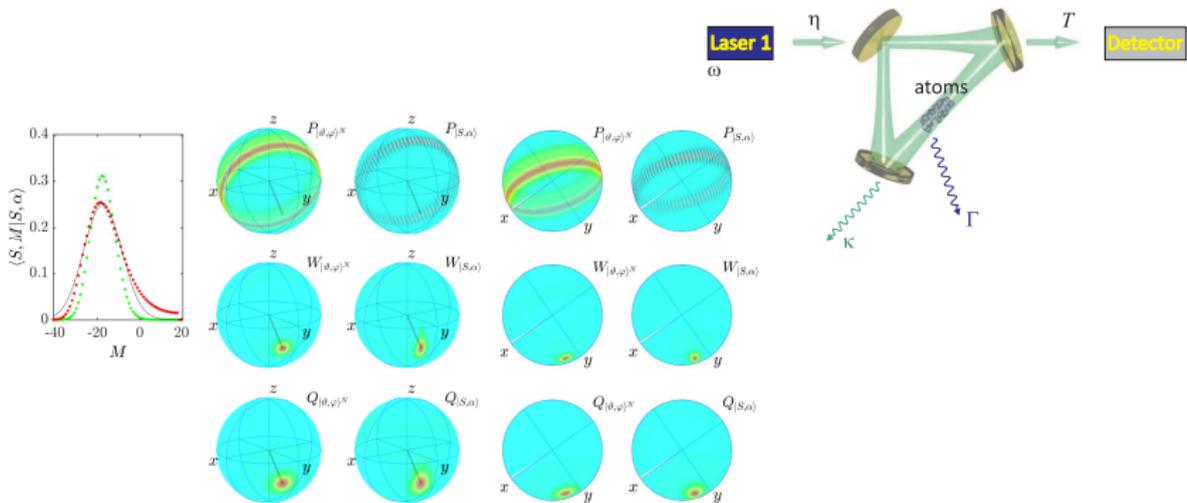
[Somech, Leppen, Shahmoon, et al., PRA **108**, 0203725 (2023) & PRX Quantum **5**, 010349 (2024) & arXiv:2404.02134]

[Song, Rey, Thompson et al., Science Adv. **11**, eadu5799 (2025)]

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$$\hat{S}_- |ss\rangle \simeq \alpha |s\rangle \quad \& \quad [\hat{S}_-, \hat{H}_{\text{eff}}] = 0 \quad \Leftrightarrow \quad 0 = \frac{d}{dt} \hat{\rho} = \frac{d}{dt} |ss\rangle \langle ss|$$



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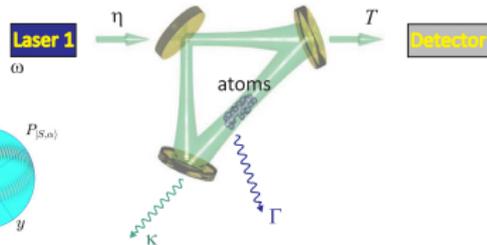
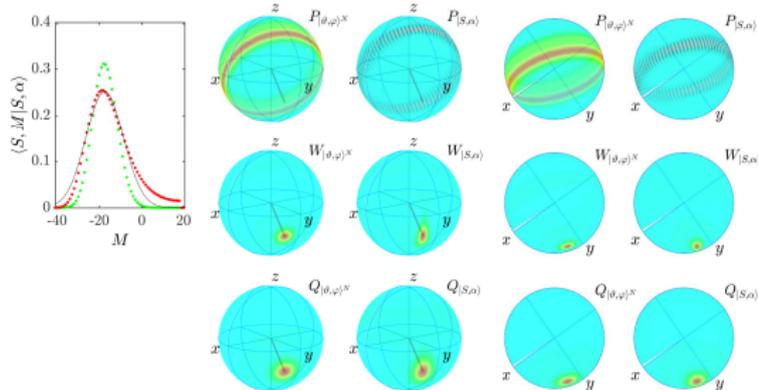
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$$\hat{S}_- |ss\rangle \simeq \alpha |s(s-1)\rangle \quad \& \quad [\hat{S}_-, \hat{H}_{\text{eff}}] = 0 \quad \Leftrightarrow \quad 0 = \frac{d}{dt} \hat{\rho} = \frac{d}{dt} |ss\rangle \langle ss|$$

P -, Wigner, and Q -function of CSS and CRSS



light scattered by a collective spin: $\hat{a}^\dagger = \hat{a}_0^\dagger + G\hat{S}_-$

[Wang, Wu et al., New J. Phys. **16**, 063039 (2014)]

[Somech, Leppenens, Shahmoon, et al., PRA **108**, 0203725 (2023) & PRX Quantum **5**, 010349 (2024) & arXiv:2404.02134]

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Spin-squeezing near a Dicke phase transition

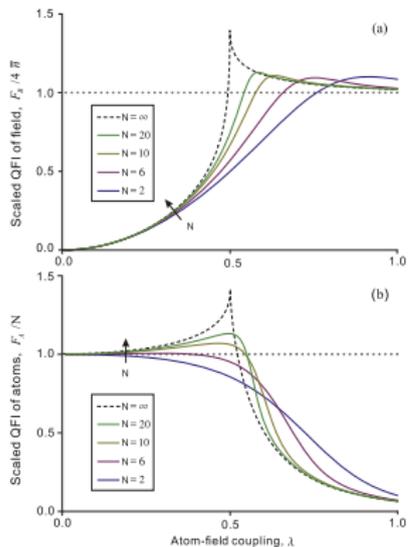


Figure 1. Scaled quantum Fisher information of the bosonic field $F_B/(4\bar{n})$ (a) and that of the atoms F_A/N (b) as a function of the coupling strength λ for a finite number of atoms $N = 2, 6, 10, 20$, as indicated by the arrow. Horizontal dotted lines: the classical (or shot-noise) limit for the field mode $F_B = 4\bar{n}$ (with mean number of bosons \bar{n}) and that of the atoms $F_A = N$. Dashed lines: analytical results of the QFI in the thermodynamic limit (i.e., $N = \infty$). For each state $\hat{\rho}_{\lambda, \bar{n}}$, the derivative of the QFI has a singularity at the critical point λ_{cr} . Other parameter: the critical coupling $\lambda_{cr} \equiv \sqrt{\omega\omega_0}/2 = 1/2$ on resonant condition $\omega = \omega_0 = 1$.

RWA, no dissipation

Quantum Fisher information and squeezing parameter

Spin-squeezing near a Dicke phase transition

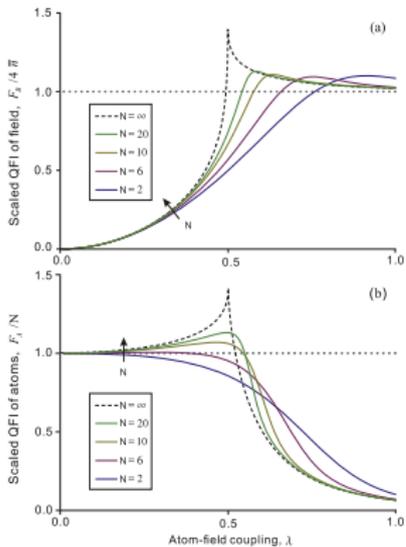


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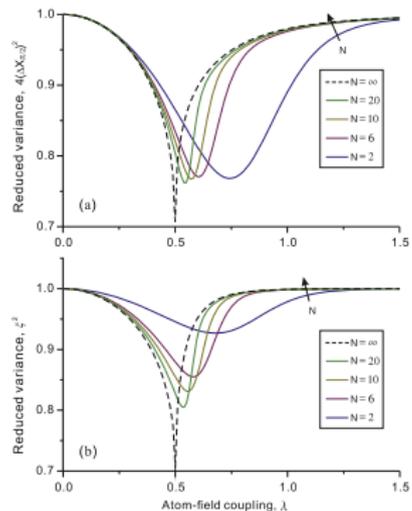


Figure 3. Degree of quadrature squeezing for the field mode $4(\Delta X_{r,2})^2$ (a), and that of spin squeezing for the atoms s^2 (b) against the coupling strength λ for the number of atoms $N = 2, 6, 10, \text{ and } 20$, as indicated by the arrow. Dashed lines: analytical results in the thermodynamic limit (i.e., $N = \infty$). The local minimum of the reduced variances indicates quadrature squeezing of $\hat{\rho}_{AB}$ at the critical point $\lambda_{cr} = 0.5$ (on resonance, as figure 1).

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Spin-squeezing near a Dicke phase transition

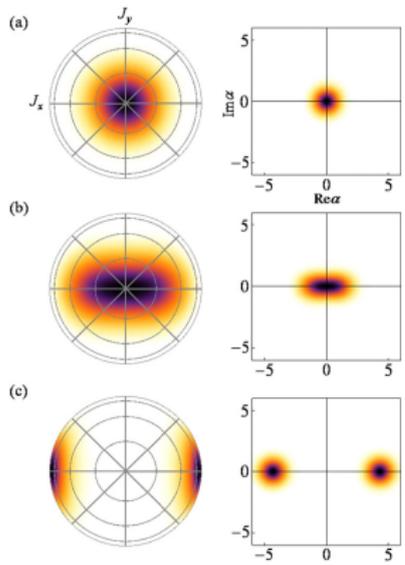


Figure 2. Quasi-probability distributions $Q_s(\theta, \phi)$ (left panel) and $Q_s(\alpha)$ (right panel) of the ground state of the Dicke Hamiltonian with $N = 20$ and the atom-field coupling strength $\lambda = 0$ (a), 0.54 (b), and 1 (c). The axes on the Bloch sphere (top view from the south pole) are given by $J_{x,z} = (\hat{J}_{x,z})$, while for that of the field mode, $\text{Re } \alpha = \langle \hat{X}_s \rangle$ and $\text{Im } \alpha = \langle \hat{X}_{s/2} \rangle$. The expectation values are taken with respect to the coherent states $|\theta, \phi\rangle$ and $|\alpha\rangle$, respectively. Other parameters: the critical coupling $\lambda_{cr} = 1/2$, the same as in figure 1. The density of Q_s is normalized by its maximal value [44, 45], i.e., $Q_{s,max} = 1$ (a), 0.557 (b), and 0.5 (c).

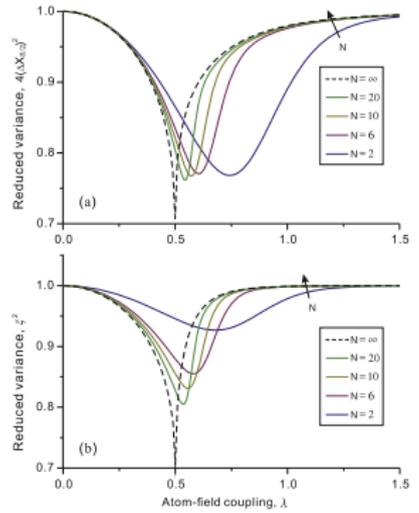


Figure 3. Degree of quadrature squeezing for the field mode $4\langle\Delta\hat{X}_{r,2}\rangle^2$ (a), and that of spin squeezing for the atoms ζ^2 (b) against the coupling strength λ for the number of atoms $N = 2, 6, 10$, and 20 , as indicated by the arrow. Dashed lines: analytical results in the thermodynamic limit (i.e., $N = \infty$). The local minimum of the reduced variances indicates quadrature squeezing of $\hat{\rho}_{AB}$ at the critical point $\lambda_{cr} = 0.5$ (on resonance, as figure 1).

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