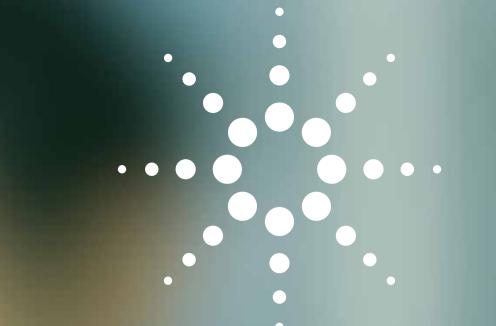


# OPTIMUM PARAMETERS FOR UV-VIS-NIR SPECTROSCOPY

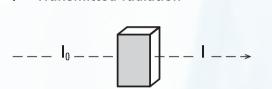


# THE BEER-LAMBERT LAW

The concentration of a solute is directly proportional to the absorbance of the solution

Where Transmittance T is defined as  $T = I/I_0$  and  $%T = 100 \times I/I_0$ 

Where  $I_0$  = Incident radiation I = Transmitted radiation



#### The Beer-Lambert Law defines Absorbance, A, as A = $-\text{Log}_{10}(I/I_0) = \text{Log}_{10}(1/T) = ECD$

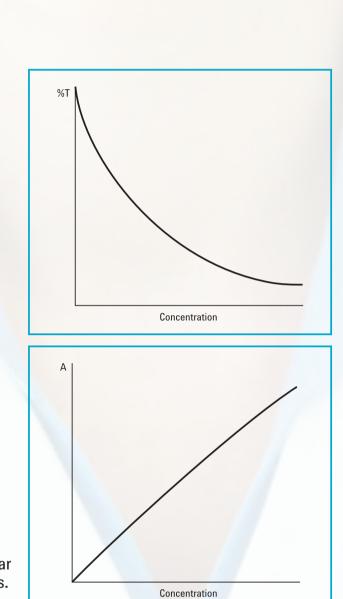
= Log (100/100 - %Absorption)

Where E = molar absorptivity, M-1cm-1 C = concentration of the solute. M D = pathlength, cm

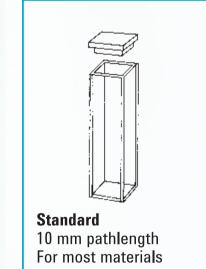
(Transmittance is often expressed as a percentage, so the relationship  $A = Log_{10}100/\%T$  is common.)

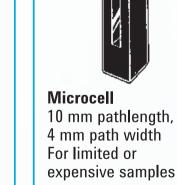
%Absorption = 100 - %TAbsorbance ≠ %Absorption

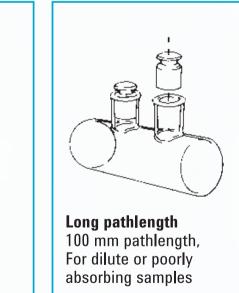
Absorbance is often more useful than Transmittance as there is a linear relationship between A and the concentration of the absorbing species. At higher concentrations there are deviations from linearity.

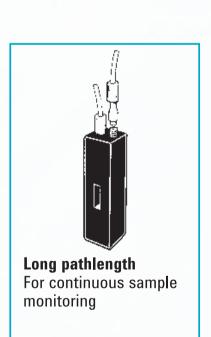


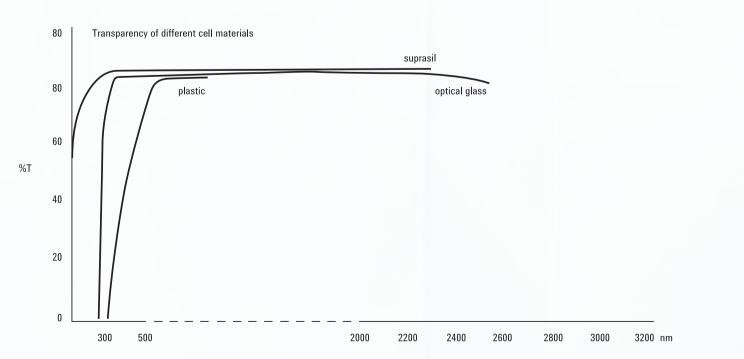
## **OPTICAL CELLS**

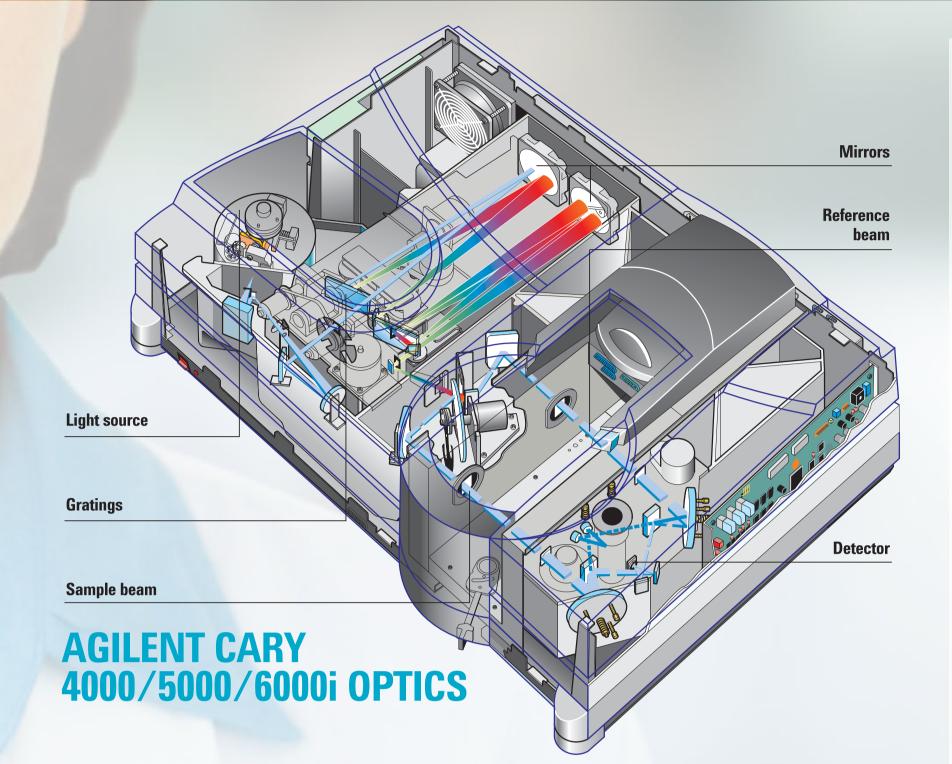


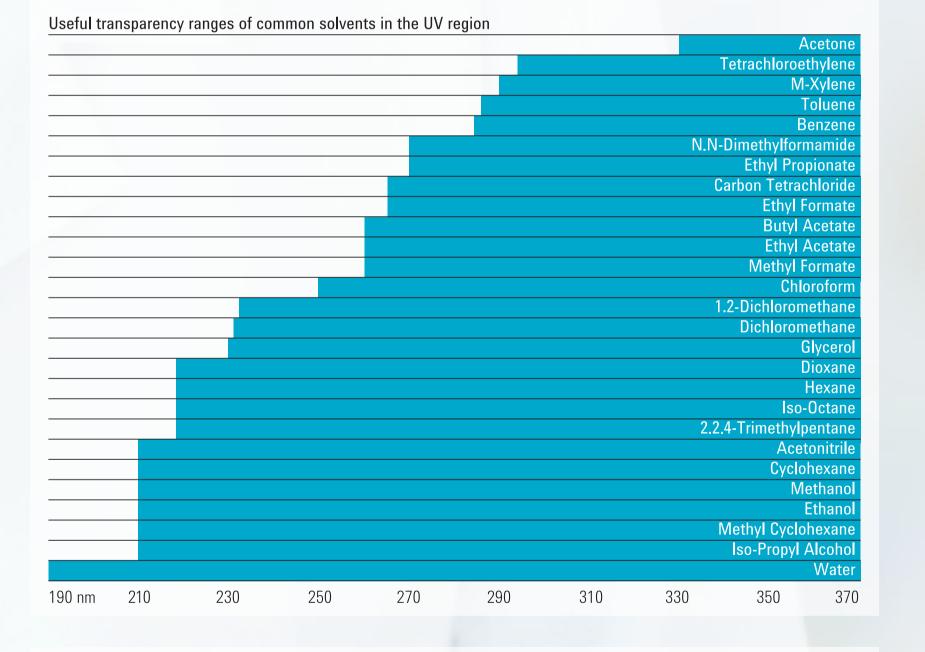












## **CONVERSION TABLES**

%T	T	А	%A	LogA	λ	cm <sup>-1</sup>
100	1	0	0	-	3300	3030
50	0.5	0.3	50	-0.52	3000	3300
10	0.1	1	90	0	2500	4000
1	0.01	2	99	0.3	2000	5000
0.1	0.001	3	99.9	0.48	1500	6666
0.01	0.0001	4	99.99	0.60	1000	10000
0.001	0.00001	5	99.999	0.70	800	12500
0.0001	0.000001	6	99.9999	0.78	600	16667
0.00001	0.000000	7	199.99999	0.85	400	25000
					200	50000
						E 74.40

## **INSTRUMENT FEATURES – WHY THEY MATTER**

#### **STRAY LIGHT**

## Wavelengths of light other than the desired wavelengths that reach

**How it is measured:** The solutions used to test stray light levels are non-transmitting at the indicated wavelengths (they do transmit at other wavelengths), so the observed transmittance is due only to stray light. The effects: The stray light level determines the maximum absorbance the instrument is able to measure. Stray light also causes deviations from the Beer-Lambert Law.

#### The 'noise' level is an indication of how stable a reading is. It determines the precision of the measurement and the detection limit of the instrument.

**How it is measured:** Noise is measured at a particular wavelength and Absorbance level. It is either specified as the peak-to-peak value (maximum deviation on either side of the Abs value) or the RMS (root mean square) which is approximately one fifth of the peak-to-peak value.

**The effects:** Poor signal-to-noise performance makes it very difficult to tell what the real Abs value is as it fluctuates. It introduces errors into both quantative and qualitative spectroscopy.

#### **WAVELENGTH ACCURACY**

#### How close the indicated wavelength is to the actual wavelength. **How it is measured:** By scanning the D<sub>2</sub> emission lines which are known to occur at 486.0 nm and 656.1 nm and observing the wavelengths at which these lines are measured.

**The effects:** Poor signal-to-noise performance makes it very difficult to tell what the real Abs value is as it fluctuates. It introduces errors into both quantative and qualitative spectroscopy.

#### **WAVELENGTH REPEATABILITY**

## The ability of the instrument to correctly return to the set wavelength

**How it is measured:** By repeatedly scanning a filter peak. If the wavelength repeatability is good, exactly the same trace should be obtained each time. **The effects:** Poor wavelength repeatability will result in errors in quantitative analysis due to wavelength shifts.

## LONG-TERM STABILITY

## The drift of the reading per unit time.

**How it is measured:** By recording the Abs of a sample at one wavelength for several hours. The specification indicates the maximum deviation from the correct Abs value per unit time.

**The effects:** Poor stability introduces errors in time-based measurements. The instrument cannot distinguish between slope due to reaction or slope introduced by the instrument. Baseline correction can be invalid.

## PHOTOMETRIC LINEARITY

#### How accurately the instrument measures absorbance with increasing concentration.

**How it is measured:** By measuring the absorbance of successive K<sub>2</sub>Cr<sub>2</sub>O<sub>7</sub> Solutions of increasing concentrations, or by measuring the additive absorption of a series of filters.

**The effects:** Poor photometric linearity will produce incorrect results.

## PHOTOMETRIC ACCURACY

## How accurately the instrument measures.

**How it is measured:** Photometric accuracy has traditionally been measured by the following methods:

Chemical standards, such as K<sub>2</sub>Cr<sub>2</sub>O<sub>7</sub> (accuracy ±0.005 Abs) Calibrated neutral density filters (accuracy is within 0.5 - 1% of their stated values or 0.002-0.004 Abs). The Double Aperture method is by far the most accurate method. It is used by major national standards laboratories to measure the absolute accuracy of their reference spectrophotometers. This method has no limitations on SBW, wavelength or temperature and can yield precisions about 2 orders better than the previously mentioned uncertainties.

The effects: Poor photometric accuracy will produce incorrect results.

# OPTIMUM SBW FOR TYPICAL COMPOUNDS **Representative Compound**

Amino Acids			
tyrptophan	279	45	4.5
tyrosine	275, 195	40, 10	4.0, 1.0
phenylalanine	258	2.2	0.2
Nucleotides			
adenosine	260	28	2.8
thymine	265	30	3
Proteins			
cytochrome c, oxidized	410	25	2.5
rhodopsin	500, 278	~90, 25	9, 2.5
ribonuclease	278	~20	2
Pigments and Dyes			
ß-carotene	480	35	3.5
chlorophyll a	660	20	2
Coenzymes			
Nicotinamide adenine dinucleotide	260	35	3.5
NADH	340, 260	50, 25	5, 2.5
Simple Organics			
benzene, vapor	253	<<0.1	<<0.01
benzene, solution	253	2	0.2
anthracene	375	3	0.3

## **INSTRUMENT PARAMETERS**

Scan was collected under

Sample absorbing 3 Abs2 nm SBW

proprocessional influence of the contract of t

0.000 1.000 2.000 3.000 4.000 min

654.9 655.3 655.7 656.1 656.5 656.9 nm

8 hours of successive scanning of a filter

443.5 444.5 445.5 446.5 447.5 448.5 nm

600.000

400.000

340 nm 2 seconds SAT 10 hours

0.0 8.9 17.8 26.7 35.6 44.5 Cond

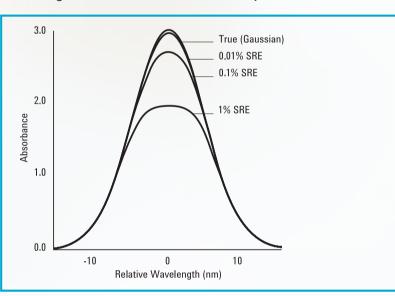
 $E(.5, .5) = \triangle T(.5)$ 

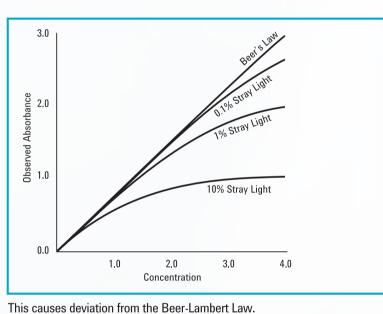
Transmittance scale error at 50%T

200.000

540 nm1.6 Abs rare bean attenuation1 second SAT

Stray light or stray radiant energy (SRE) is defined as the percentage of radiation reaching the detector whose wavelengths are outside the selected spectral band.





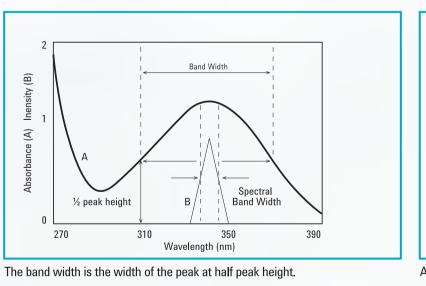
The effects of SRE are: decreased absorbance readings changes in band shape.

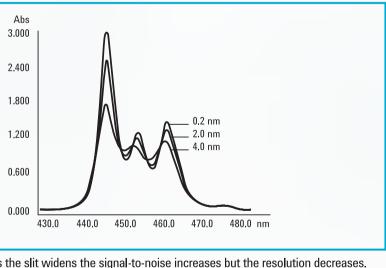
The level of SRE determines the maximum Abs measurable by the instrument.

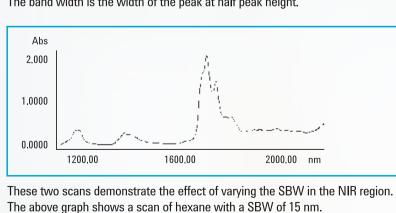
## **RESOLUTION**

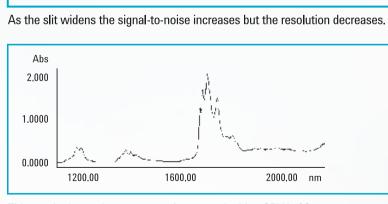
**Spectral Band Width** should be set to one-tenth of the band width.

**Resolution** the spectrophotometer's ability to distinguish between two absorbance bands which are close together. **Data Interval** a minimum ratio of SBW: Data Interval of 3:1 should be set to ensure that no spectral detail is lost.









This graph shows the same sample scanned with a SBW of 3 nm –

with considerably better resolution.